

Recent results  
from  
Atmospheric  $\nu$  and K2K

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# Super- Kamiokande and K2K experiment

World Largest  
Water Cherenkov detector

● 1996.4 Start data taking

SK-I  
1000 m underground  
50,000 ton  
(22,500 ton fid.)  
11,146 20 inch PMTs  
1,885 anti-counter PMTs

● 1999.3 K2K start

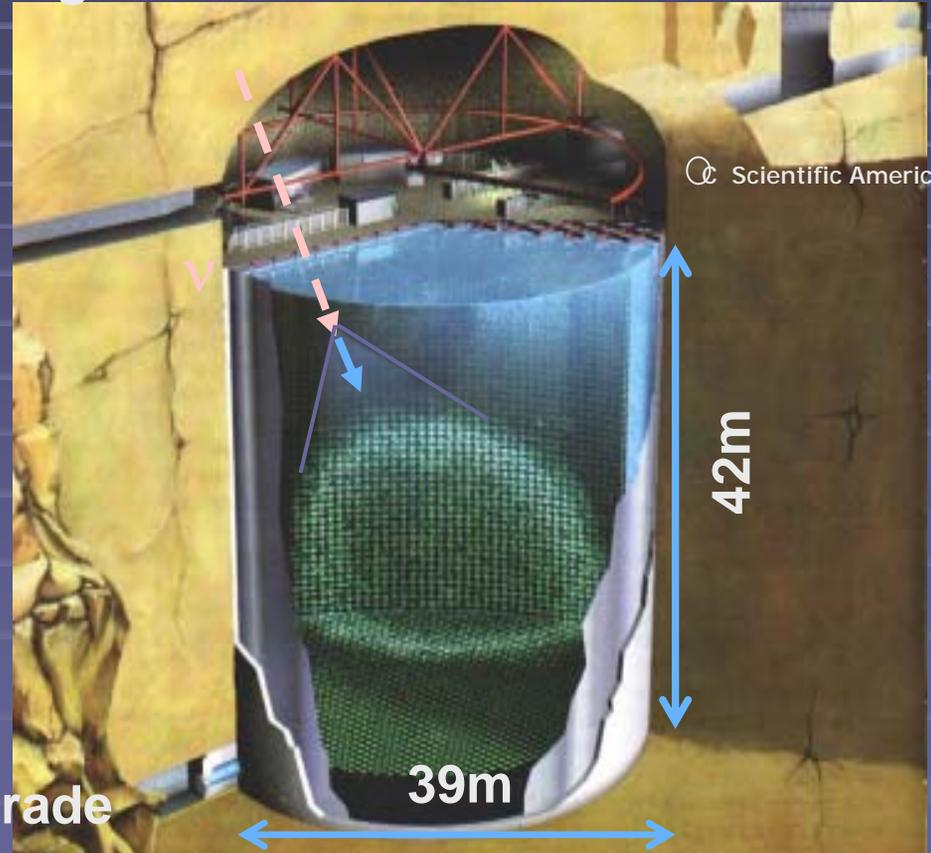
● 2001.7 Detector upgrade

● 2001.11 Accident

partial reconstruction of the detector

● 2002.10 resume data taking

● 2002.12 resume K2K beam



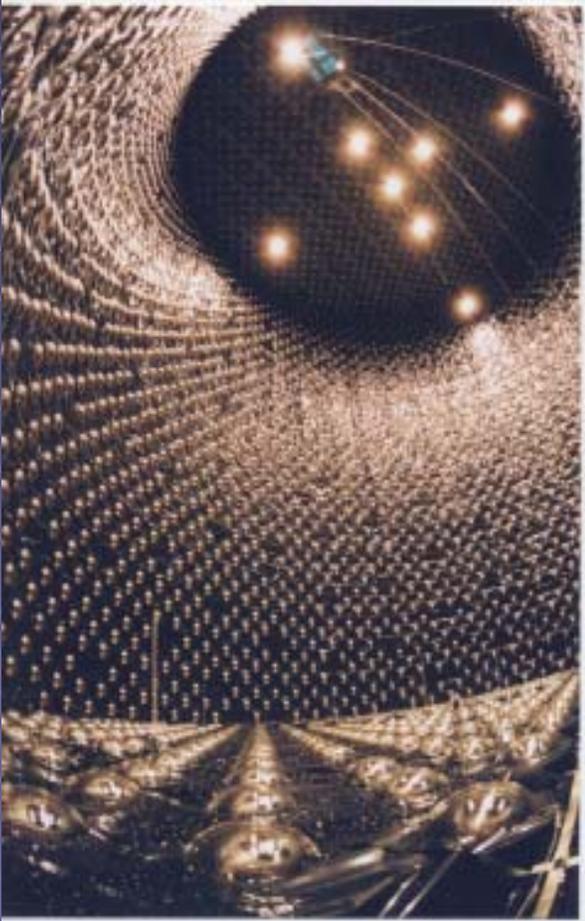
K2K-I

SK-I

SK-II  
K2K-II

# SK is back !

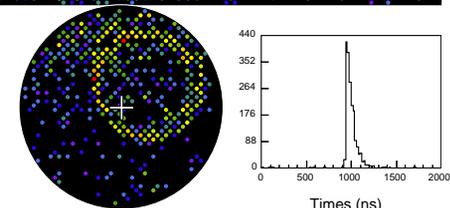
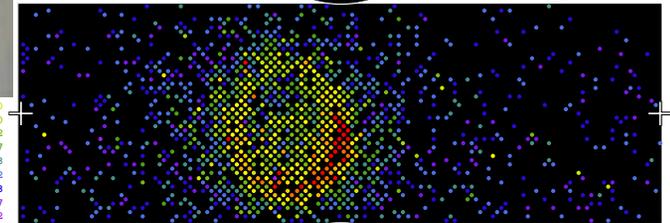
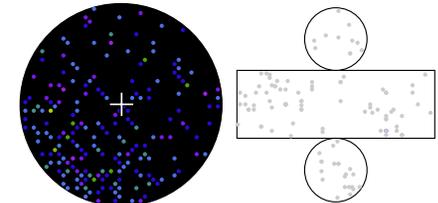
Full water on 10-Dec.-2002



Jan.-2003, fully contained event

Super-Kamiokande

Event 5348354  
3:35  
Time, 8472 pE  
0 pE (in-time)  
03  
cm  
d



Sep.-2002, before water filling

# Outline of this talk

- Atmospheric results from SK-I
  - Entire re-analysis with new  $\nu$ -flux,  $\nu$ -int model.
- K2K-I results
  - $\nu_{\mu}$  disappearance
  - $\nu_e$  appearance search
  - Study  $\nu$  interaction
- Status of SK-II / K2K-II

# Neutrino oscillation

Two neutrino case

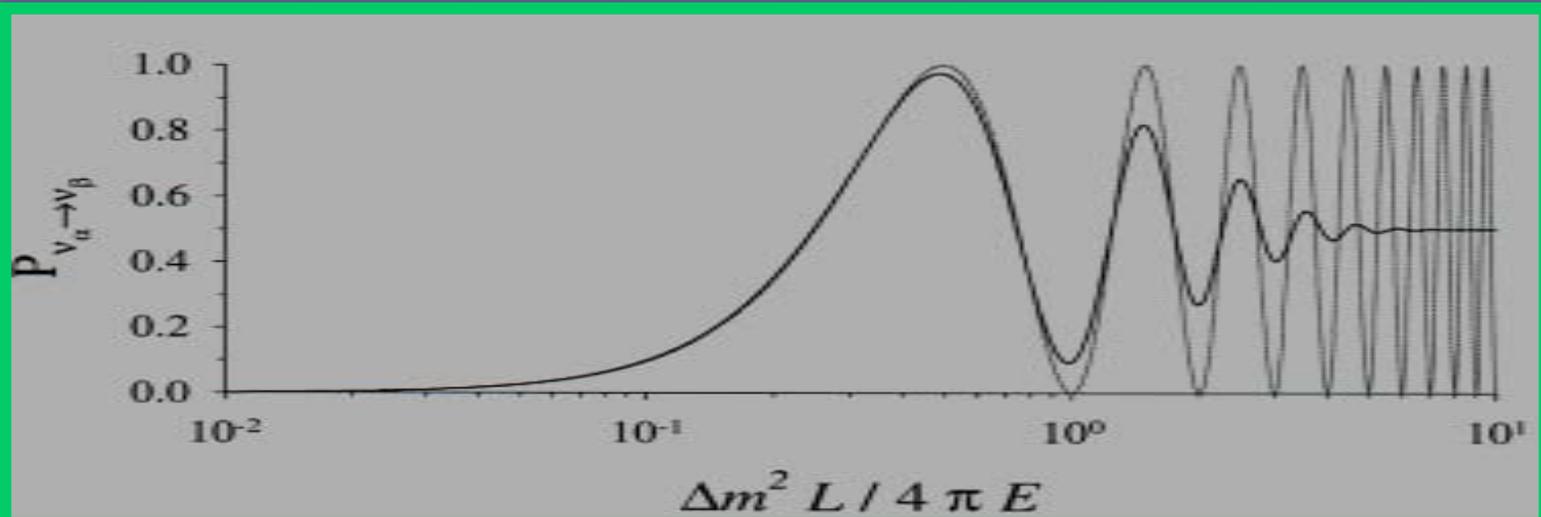
$$\begin{pmatrix} \nu_\alpha \\ \nu_\beta \end{pmatrix} = \begin{pmatrix} \cos\theta & \sin\theta \\ -\sin\theta & \cos\theta \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \end{pmatrix}$$

$$P(\nu_\alpha \rightarrow \nu_\beta) = \sin^2 2\theta \sin^2(1.27 \Delta m^2 L / E)$$

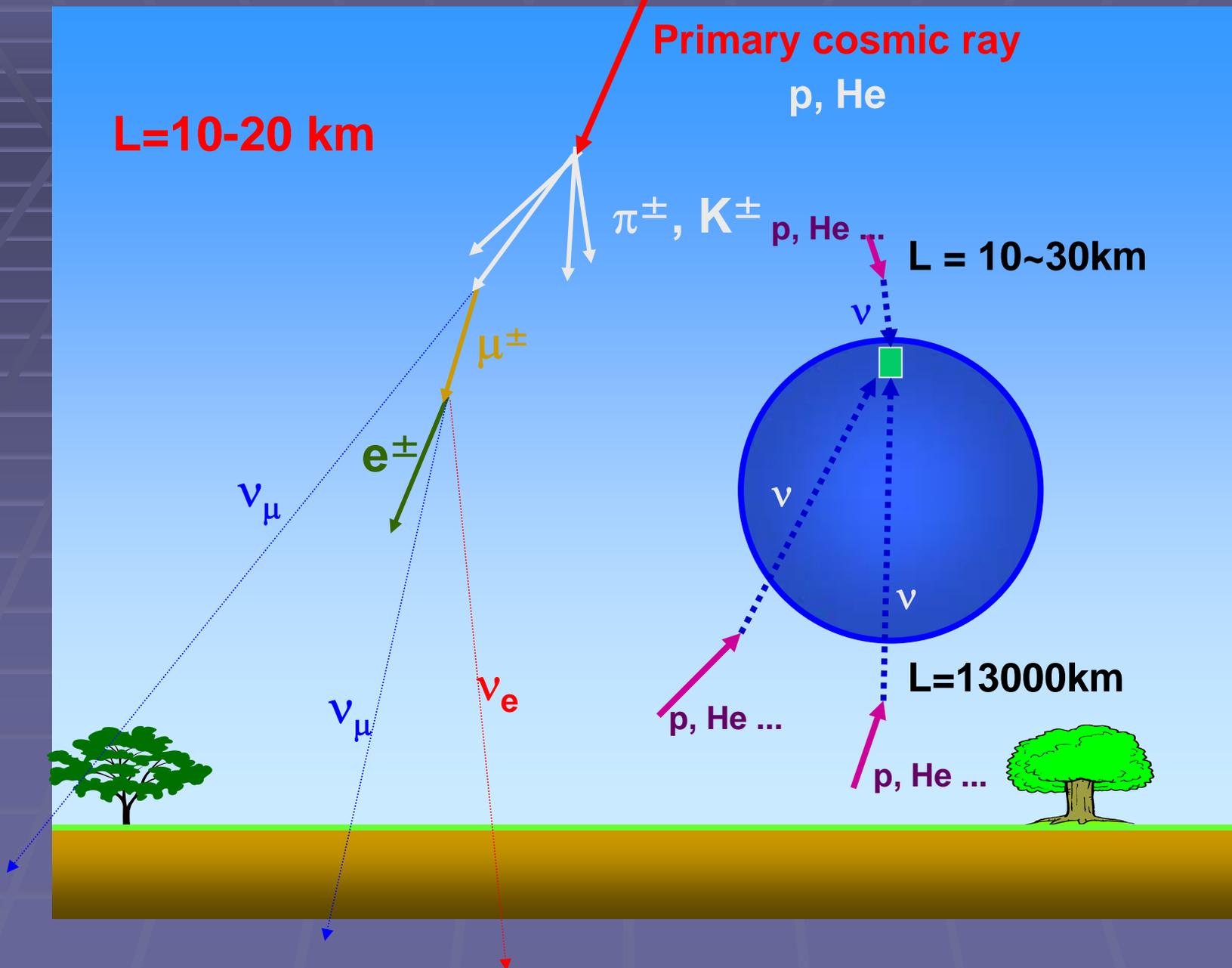
$$\Delta m^2 = m_2^2 - m_1^2 \text{ (eV}^2\text{)}$$

L (km): Distance from source to detector

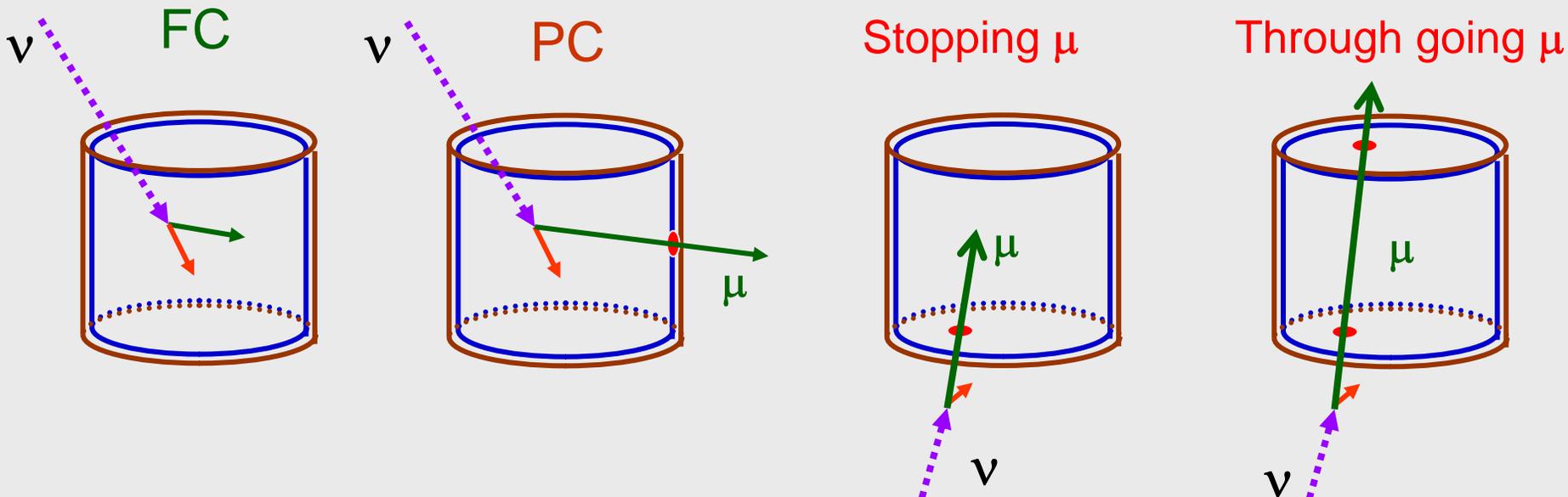
E (GeV): Neutrino energy



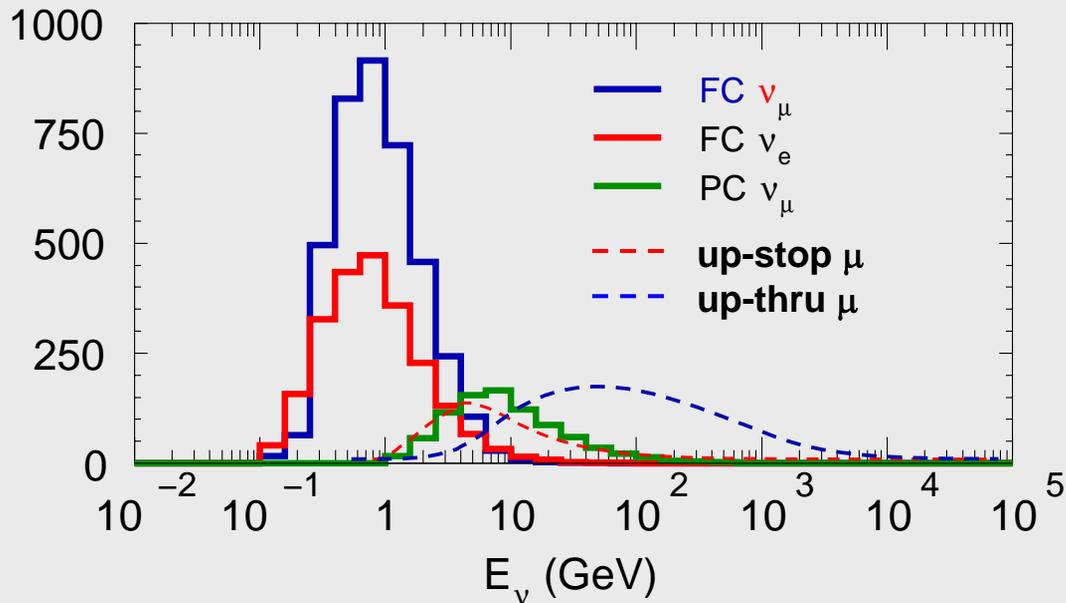
# Atmospheric neutrinos



# Atmospheric $\nu$ categories at SK



Energy spectrum of  $\nu$  for each event category



# Summary of SK-I contained events

Preliminary!

## Sub-GeV (Fully Contained)

$E_{vis} < 1.33 \text{ GeV}$ ,  
 $P_e > 100 \text{ MeV}$ ,  $P_\mu > 200 \text{ MeV}$

	Data	MC(Honda)
1ring e-like	3353	3013.9
μ-like	3227	4466.9
Multi ring	2361	2959.0
(μ-like)	(208)	(346.4)
<b>Total</b>	<b>8941</b>	<b>10439.8</b>

$$\frac{(\mu/e)_{data}}{(\mu/e)_{MC}} = 0.649^{+0.016}_{-0.016} \pm 0.051$$

## Multi-GeV

Fully Contained ( $E_{vis} > 1.3 \text{ GeV}$ )

	Data	MC(Honda)
1ring e-like	746	700.4
μ-like	651	948.2
Multi ring	1504	1944.6
(μ-like)	(439)	(739.4)
<b>Total</b>	<b>2901</b>	<b>3593.2</b>

Partially Contained (assigned as μ-like)

<b>Total</b>	<b>913</b>	<b>1149.8</b>
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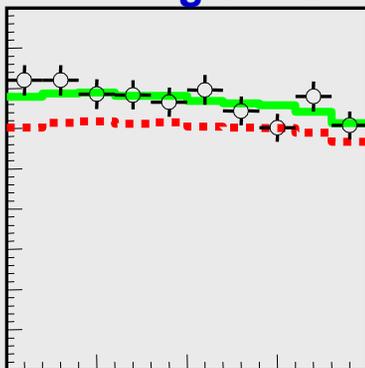
$$\frac{(\mu/e)_{data}}{(\mu/e)_{MC}} = 0.700^{+0.032}_{-0.030} \pm 0.083$$

# Atmospheric $\nu$ zenith angle distribution

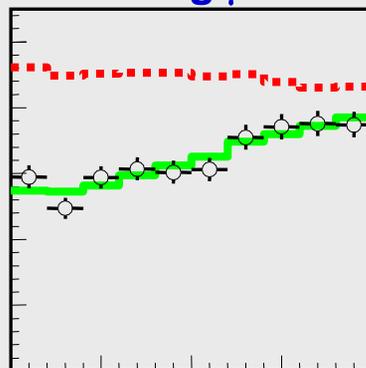
..... Honda

— Best fit ( $\sin^2 2\theta = 1.0, \Delta m^2 = 2.0 \times 10^{-3} \text{ eV}^2$ )

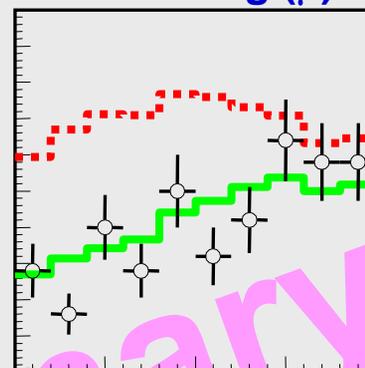
Sub GeV  
1ring e-like



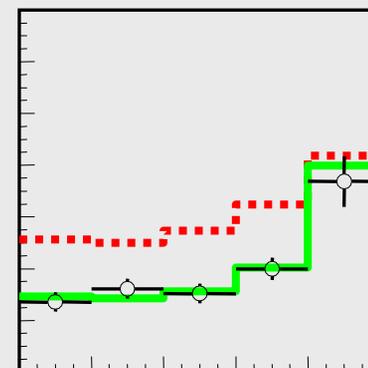
Sub GeV  
1ring  $\mu$ -like



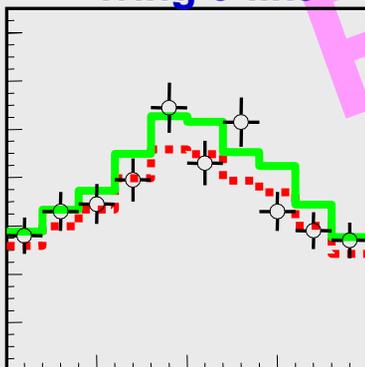
Sub GeV  
Multi ring ( $\mu$ )



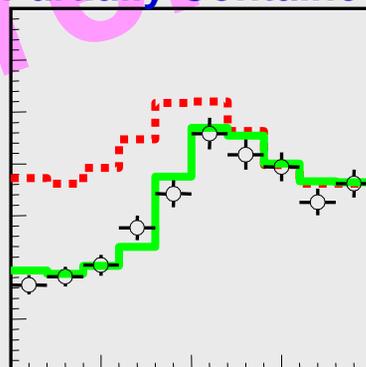
Upward stopping  $\mu$



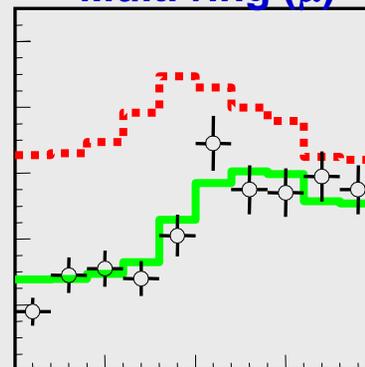
Multi GeV  
1ring e-like



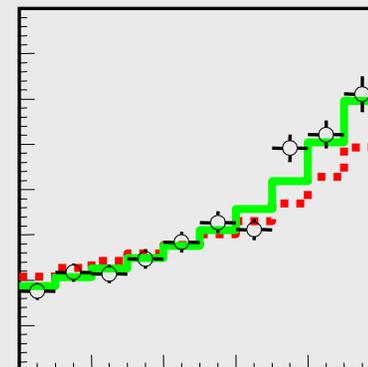
Multi-GeV 1ring  $\mu$ -like  
+ Partially Contained



Multi GeV  
Multi ring ( $\mu$ )

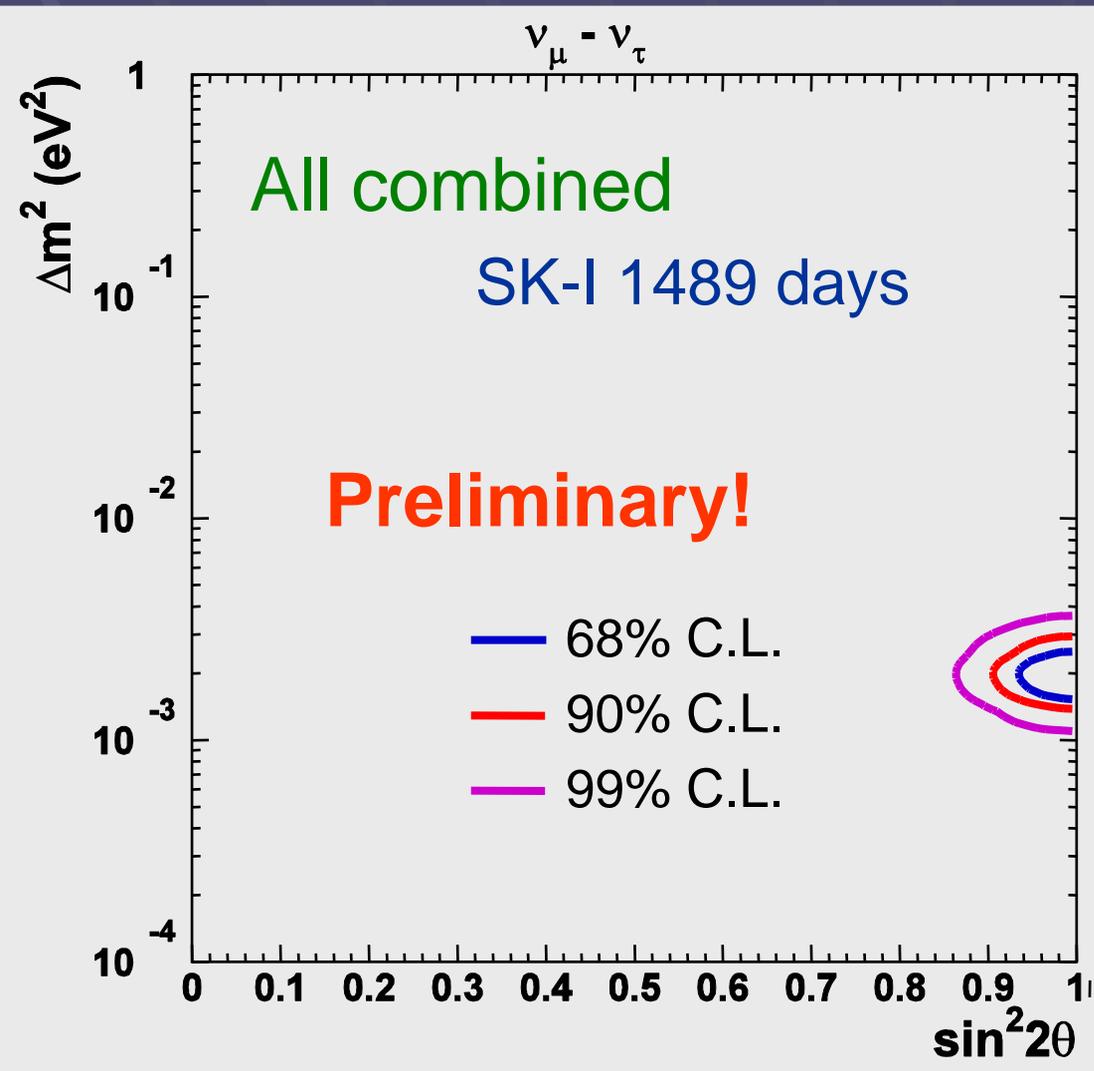


Upward  
through going  $\mu$



Preliminary!

# Allowed region of the oscillation parameters (subGeV+multiGeV+PC+MultiRing+Up $\mu$ ) (complete SK-I data set)



Assuming  $\nu_\mu \leftrightarrow \nu_\tau$  oscillation

Best fit

$$\chi^2_{\min} = 170.8/170 \text{ d.o.f.}$$

at ( $\sin^2 2\theta$ ,  $\Delta m^2$ )

$$= (1.0, 2.0 \times 10^{-3} \text{ eV}^2)$$

90% confidence level  
allowed region

$$\sin^2 2\theta > 0.9$$

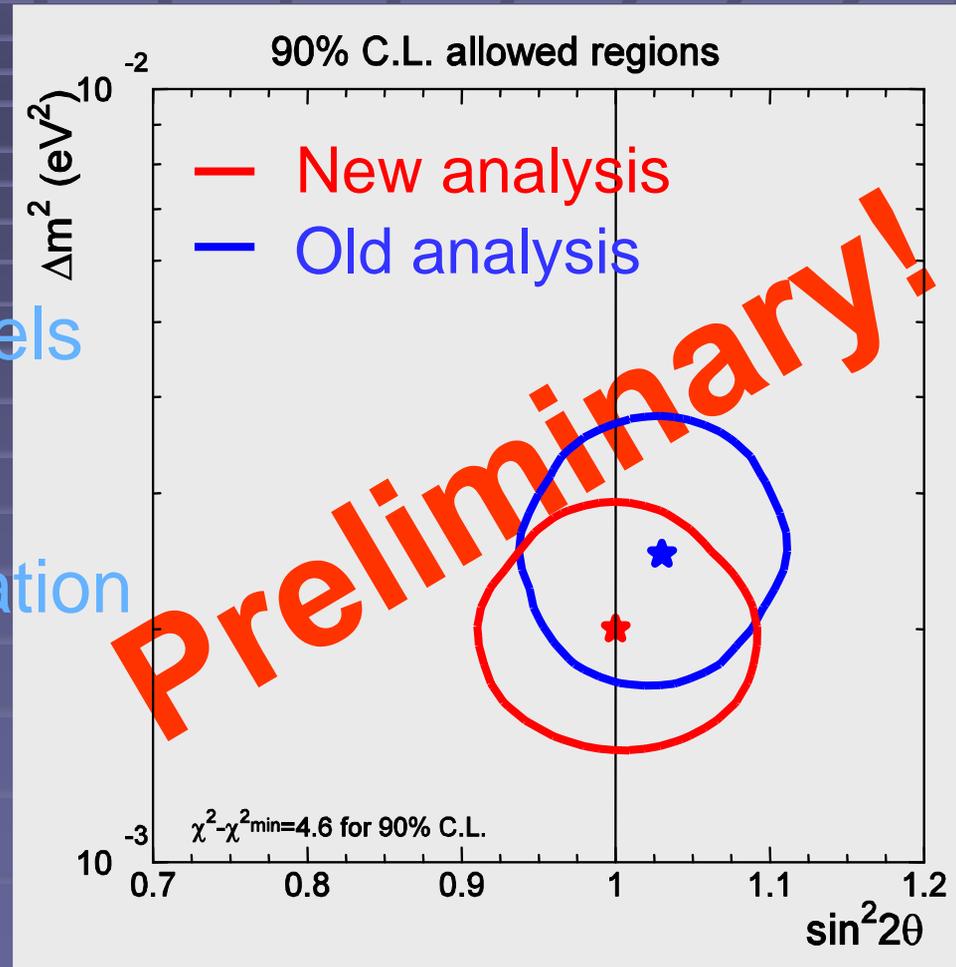
$$1.3 \times 10^{-3} < \Delta m^2 < 3.0 \times 10^{-3} \text{ (eV}^2)$$

Assuming null oscillation

$$\chi^2 = 445.2/172 \text{ d.o.f.}$$

# Comparison between old and new results

- Neutrino flux  
(Honda 1995 → Honda 2001)
- Neutrino interaction models  
(several improvements,  
agree with K2K near data)
- Improved detector simulation
- Improved event  
reconstruction tools



Each change slightly shifted  
the allowed region to lower  $\Delta m^2$

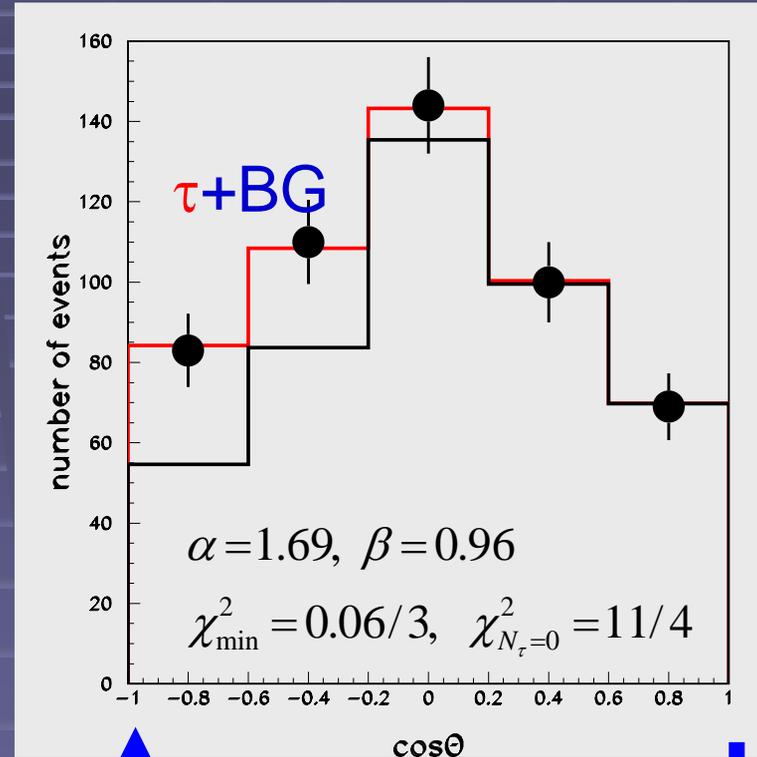
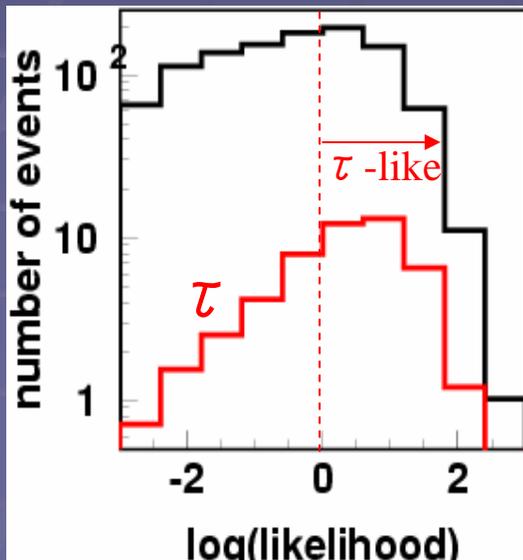
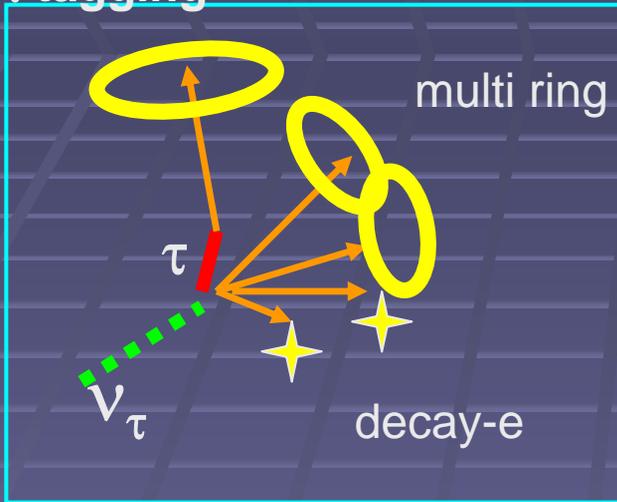
# Evidence for $\nu_{\mu}-\nu_{\tau}$

- Search for  $\nu_{\tau}$ -cc like events
  - Multi-GeV e-like multi ring events
  - $\tau$ -likelihood
- Search for mixing suppression by Matter effect
- Absolute NC rate by  $\text{NC}1\pi^0$ 
  - Use  $\pi^0$  rate measurement at K2K 1kt detector

# zenith angle dist. of cc $\tau$ -enhanced sample

$\tau$ -like selection;  $\text{eff}_\tau=44\%$ ,  $\text{S/N}=8\%$

cc- $\tau$  tagging



Consistent with  $\nu_\mu \leftrightarrow \nu_\tau$



$$N_\tau = 145 \pm 44(\text{stat})_{-16}^{+11}(\text{sys})$$

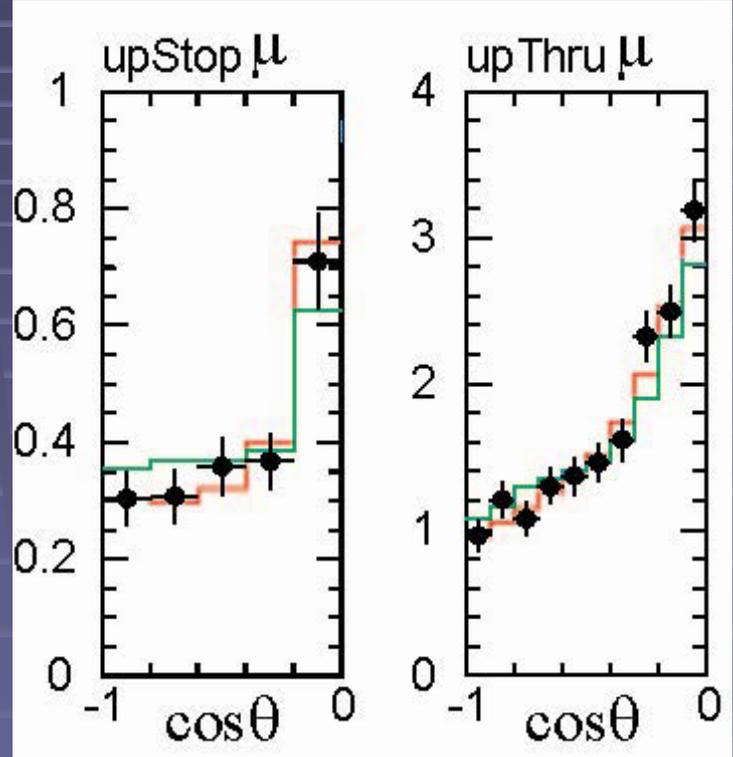
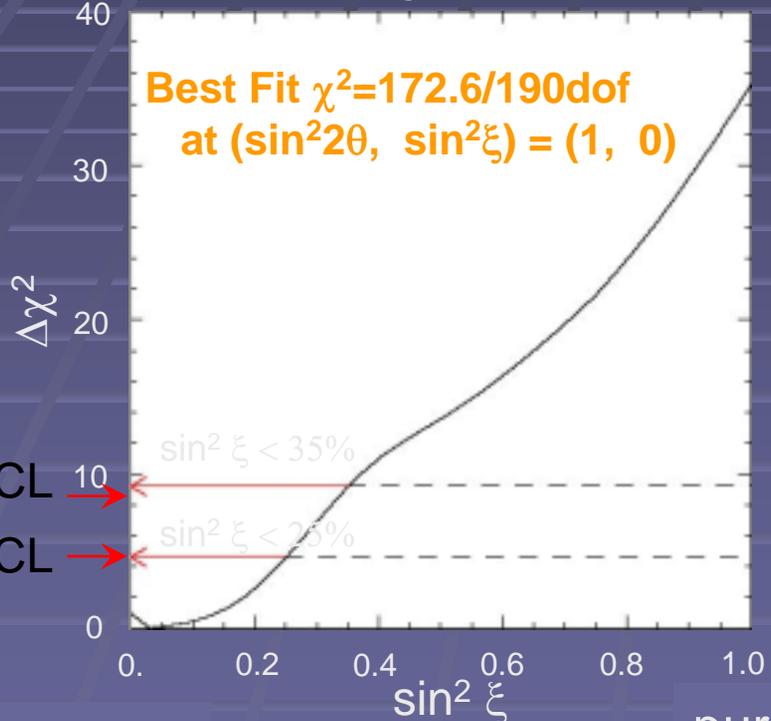
(86 expected)

# limit on $\nu_\mu \Leftrightarrow \nu_s$ admixture ( 4-flavor mixing)

(Following Fogli, Lisi, and Morrone, hep-ph/000299)

$$\nu_{\mu \rightarrow} \quad (\cos \xi \nu_\tau + \sin \xi \nu_s)$$

$\chi^2$  scan along  $\Delta m^2 = 3.2 \times 10^{-3} \text{eV}^2$



pure  $\nu_\mu - \nu_\tau$

pure  $\nu_\mu - \nu_s$

— pure  $\nu_\mu - \nu_\tau$

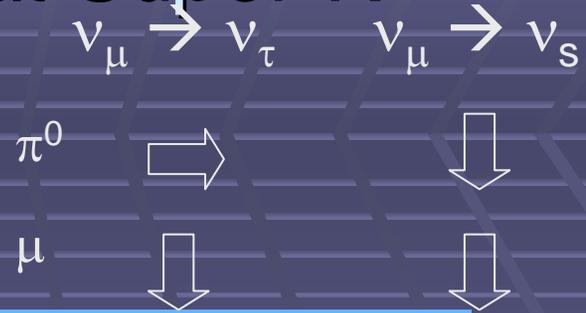
— pure  $\nu_\mu - \nu_s$

pure  $\nu_\mu - \nu_s$  is excluded by  $\Delta\chi^2 > 30$  !

# $\pi^0$ in atm- $\nu$ sample at Super-K

Check

$\nu_\mu \leftrightarrow \nu_\tau$  and  $\nu_\mu \leftrightarrow \nu_s$  hypotheses  
by a **NC rate** measurement



SK Data set : 1489 days

(\*) normalized by livetime

SK-atm $\nu$	Data	MC(*)
$\pi^0$	475	483.8
1-R FC $\mu$	3878	5415.1
$\pi^0/\mu$	<b>0.122</b> $\pm 5\%$ $\pm 7\%$ (stat) (sys)	0.089 $\pm \sim 30\%$ <b>0.087</b> $\pm \sim 13\%$ (sys)

Use K2K results

- $(\pi^0/\mu)_{MC}$
- No osc  
0.087
- Pure  $\nu_\mu - \nu_\tau$   
0.124
- Pure  $\nu_\mu - \nu_s$   
0.103

Detector systematics

Particle ID

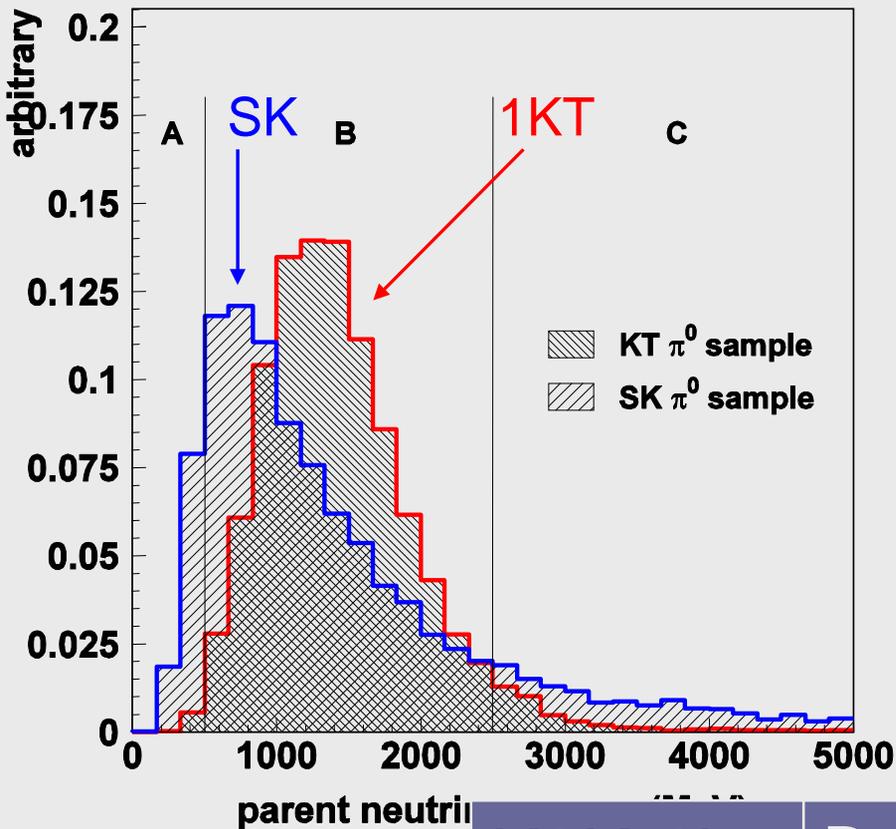
Ring counting, etc..

K2K measurement ( 8%)

SK- K2K  $\nu$ -flux difference (4%), etc..

@  $\sin^2 2\theta = 1.0$   
 $\Delta m^2 = 2.0 \times 10^{-3} \text{ eV}^2$

# NC $\pi^0$ measurement at K2K



	1kt	SK
NC frac.	86%	85%
Eff. for $\pi^0$	47%	46%

Ev producing  $\pi^0$  for 1kt and SK

K2K-1kt

Data

MC(\*)

$\pi^0$

2496

2582.3

1-R FC  $\mu$

22612

22545.2

$\pi^0/\mu$

**0.110**  $\pm 2\%$   $\pm \underline{8\%}$

**0.115**  $\pm \underline{\sim 30\%}$

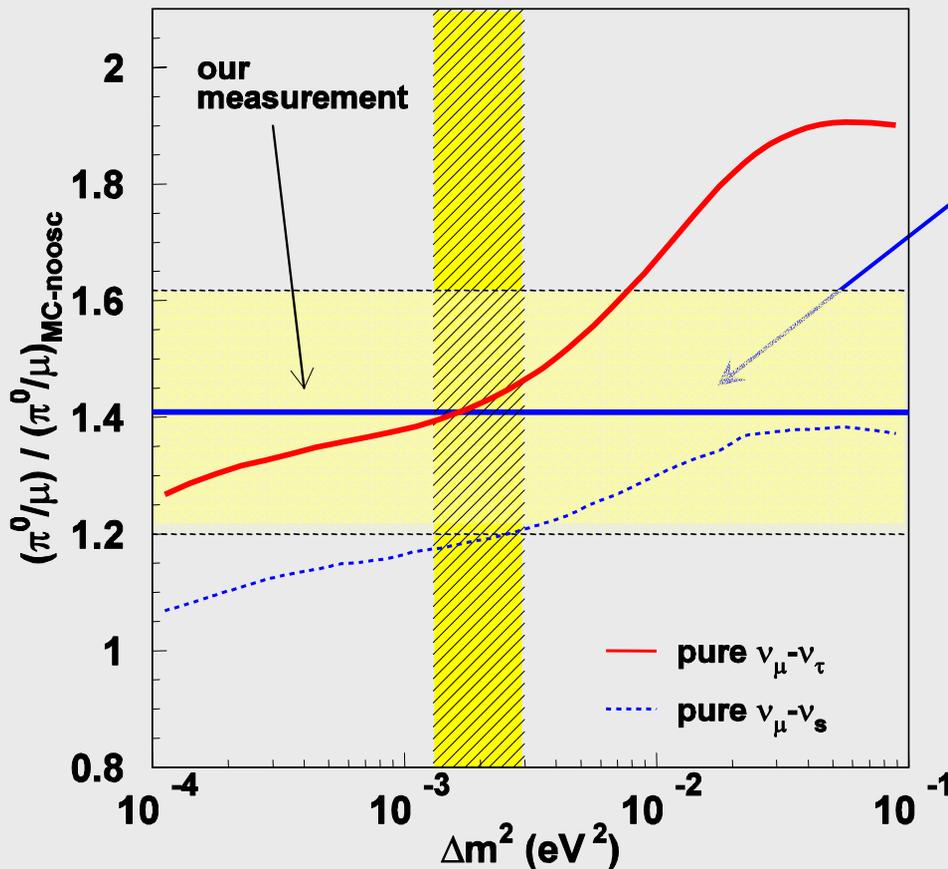
(stat)

(sys)

(sys)

# Atm- $\nu$ $\pi^0/\mu$ for data, $\nu_\mu-\nu_\tau$ and $\nu_\mu-\nu_s$ hypothesis @ $\sin^2 2\theta=1.0$ $\Delta m^2=2.0 \times 10^{-3} \text{ eV}^2$

$\nu_\mu-\nu_\tau$  is consistent



$(\pi^0/\mu) / (\pi^0/\mu)_{\text{noosc}}$	
● Data	1.41
● Pure $\nu_\mu-\nu_\tau$	1.42
● Pure $\nu_\mu-\nu_s$	1.19

# 3 Flavor Mixing

- ★ If neutrinos are massive particles, then it is possible that the **mass eigenstates** and the **weak eigenstates** are not the same:

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

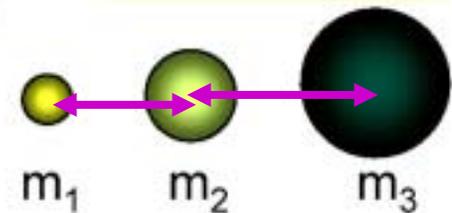
Weak eigenstates  
„flavor eigenstates“



3 independent parameters  
+ 1 complex phase

$\theta_{12}, \theta_{23}, \theta_{13}$   
+  $\delta$

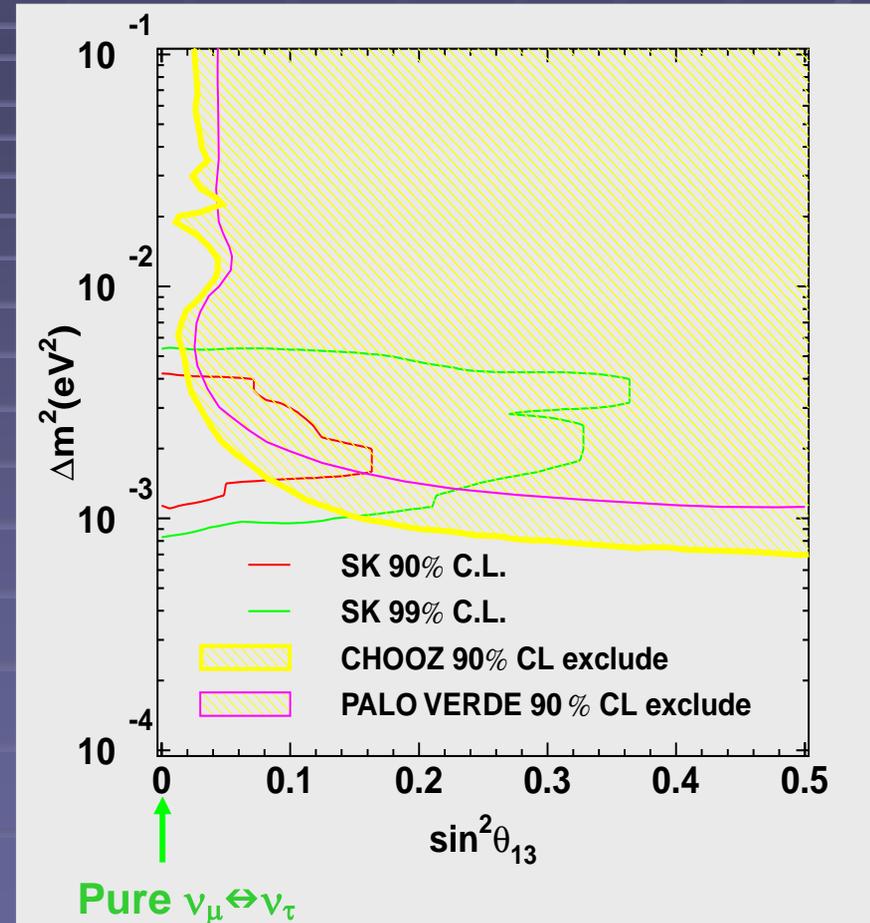
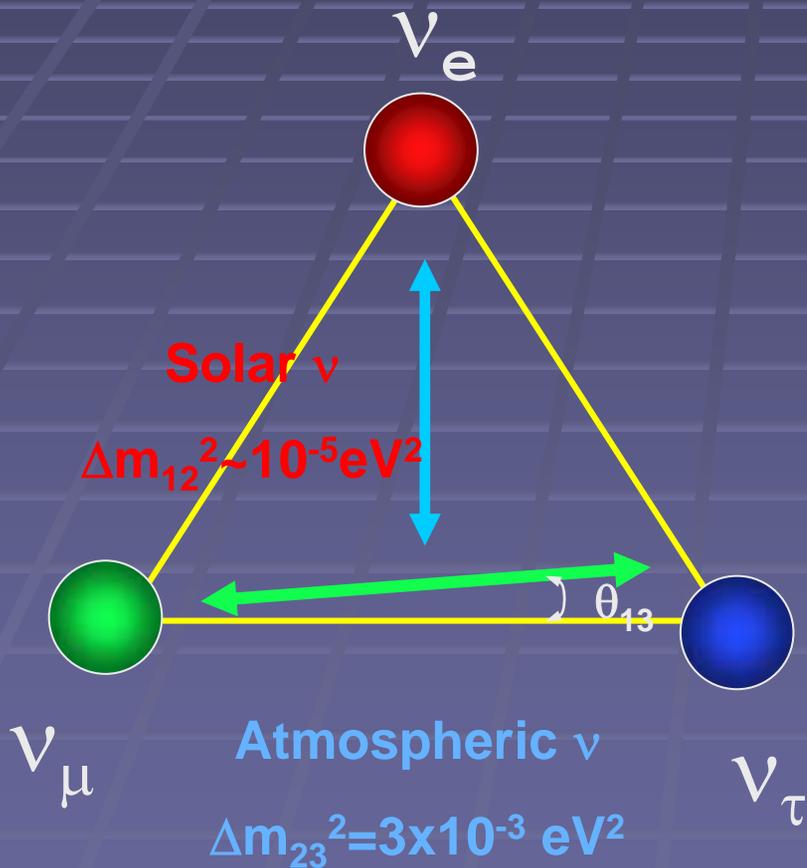
Mass eigenstates



$\Delta m^2_{12}, \Delta m^2_{23}$

MNS (Maki-Nakagawa-Sakata) matrix

# Allowed region for active 3-flavor oscillations

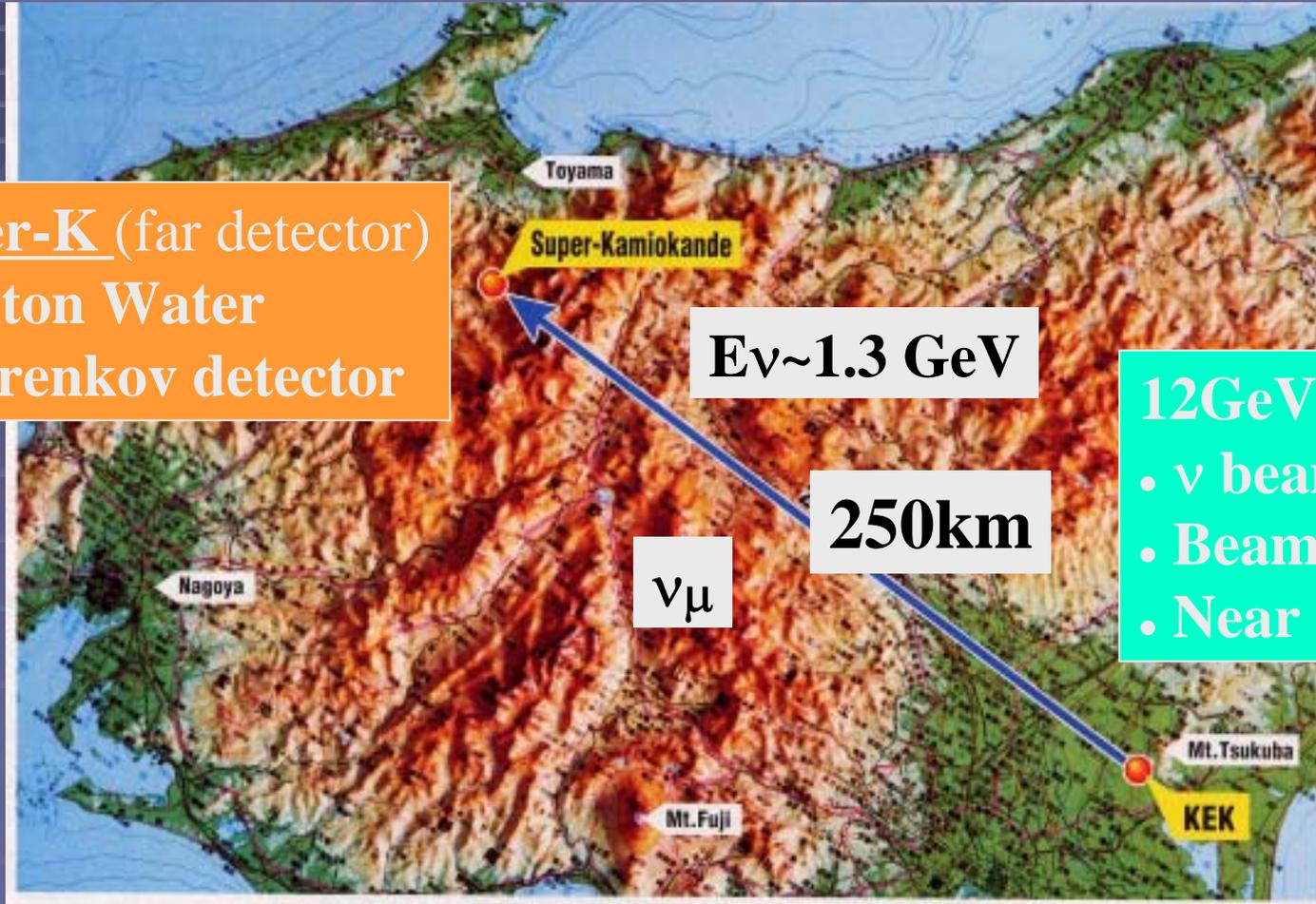


consistent with CHOOZ's excluded region

# K2K experiment

# KEK to Kamioka Neutrino Oscillation Experiment

Super-K (far detector)  
50 kton Water  
Cherenkov detector



$E_\nu \sim 1.3 \text{ GeV}$

250km

$\nu_\mu$

12GeV PS@KEK

- $\nu$  beam line
- Beam monitor
- Near detectors

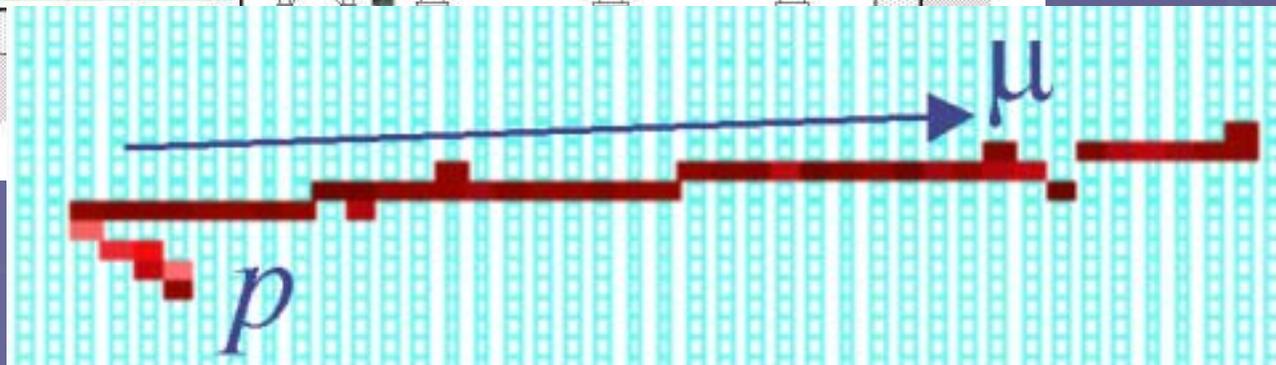
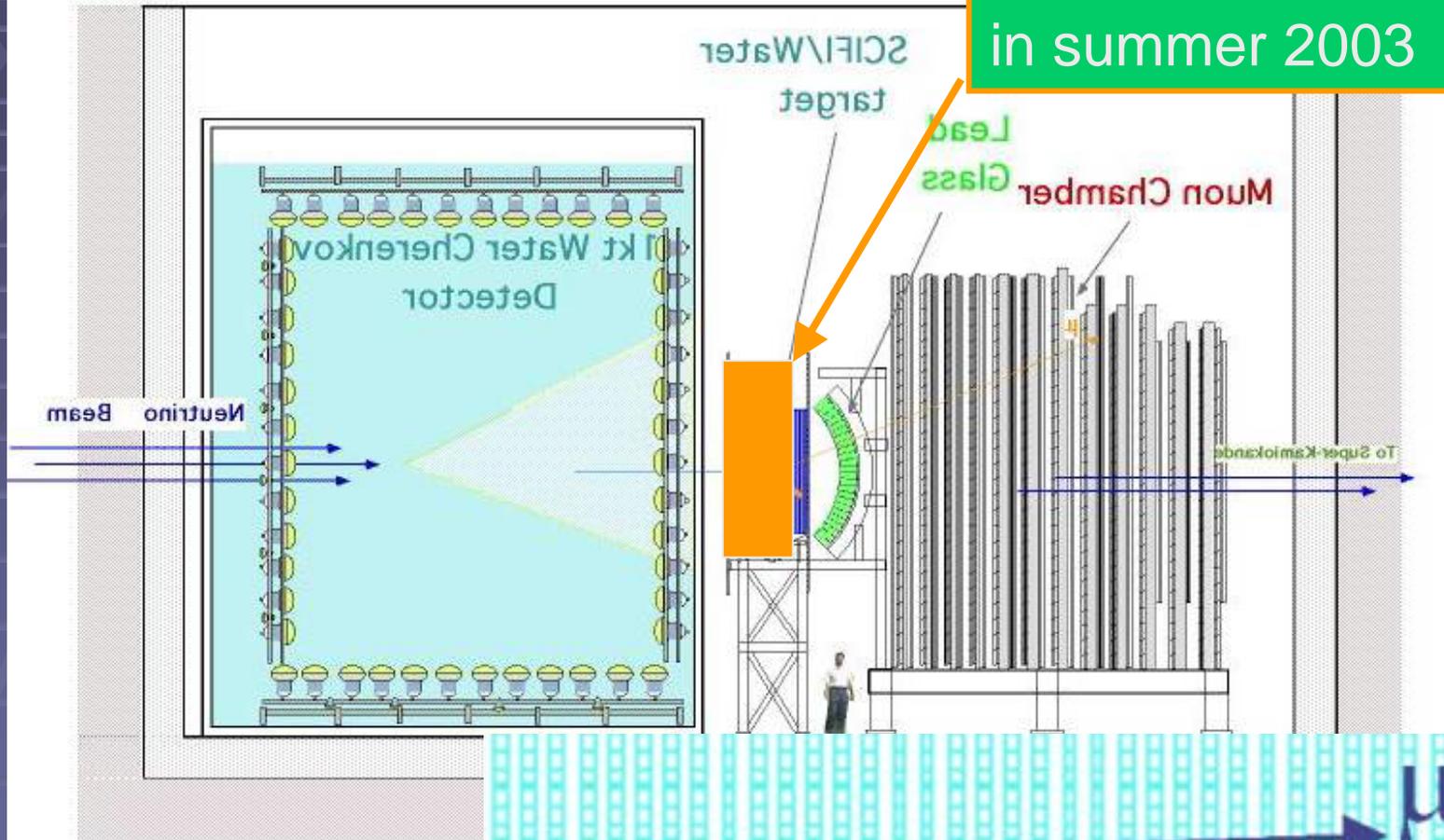
# K2K collaboration



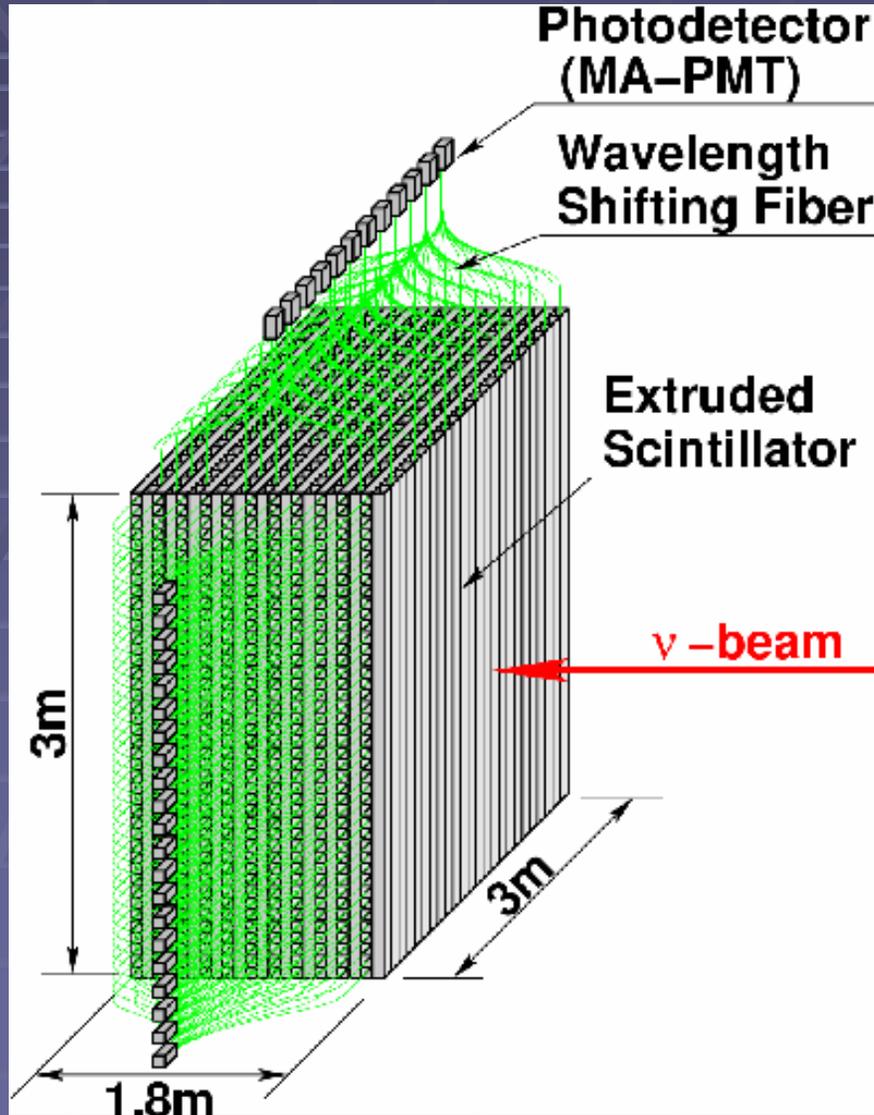
- University of Barcelona
- Boston University
- Chonnam National University
- Dongshin University
- University of Geneva
- Hiroshima University
- ICRR
- Inst. for Nuclear Research, Moscow
- KEK
- Kobe University
- Korea University
- Kyoto University
- Massachusetts Institute of Technology
- Niigata University
- Okayama University
- University of Rome
- "La Sapienza" Saclay (DSM-DAPNIA)
- Seoul National University
- SUNY at Stony Brook
- Tokyo University of Science
- Tohoku University
- University of California, Irvine
- University of Hawaii
- University of Tokyo
- University of Washington
- University of Valencia
- Warsaw University

# K2K near detector

SciBar has been installed in summer 2003



# SciBar detector



## Full active

Large Volume:

$(300 \times 300 \times 166) \text{ cm}^3 \sim 15\text{tons}$

Fine segment:  $2.5 \times 1.3 \times 300 \text{ cm}^3$

Large Light Yield:

$7 \sim 20$  photo-electrons/cm for MIP

the factor 3 of 7~20 comes from the fiber attenuation.

Particle ID:

$p/\pi$  :  $dE/dx$

$\mu/\pi$  : range

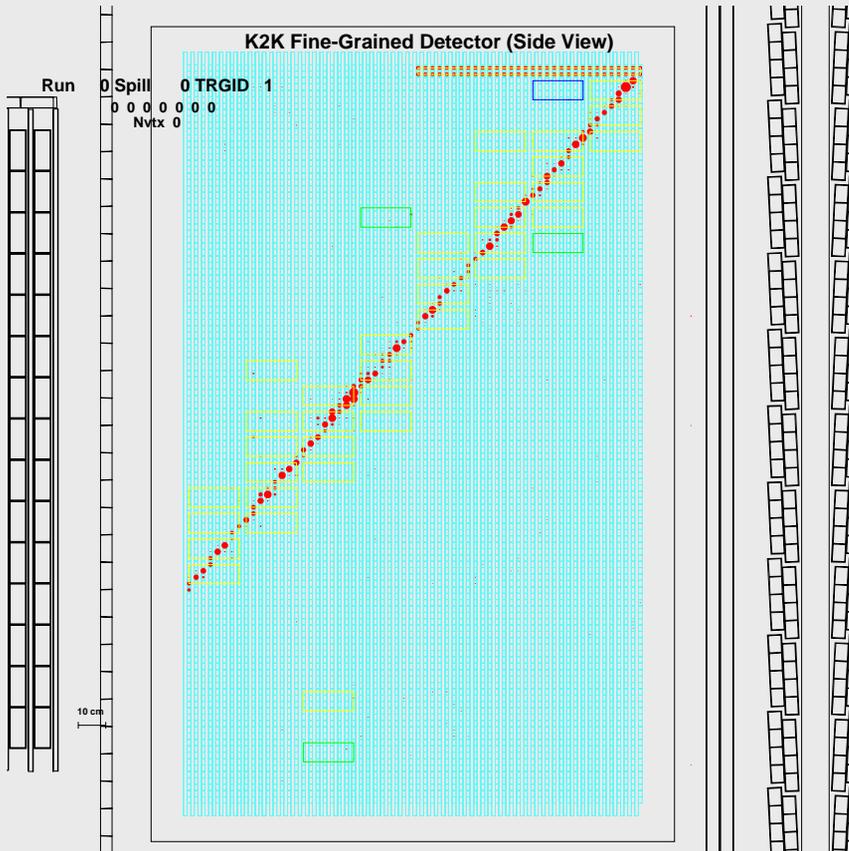
Proton Momentum:

by  $dE/dx$  and the range

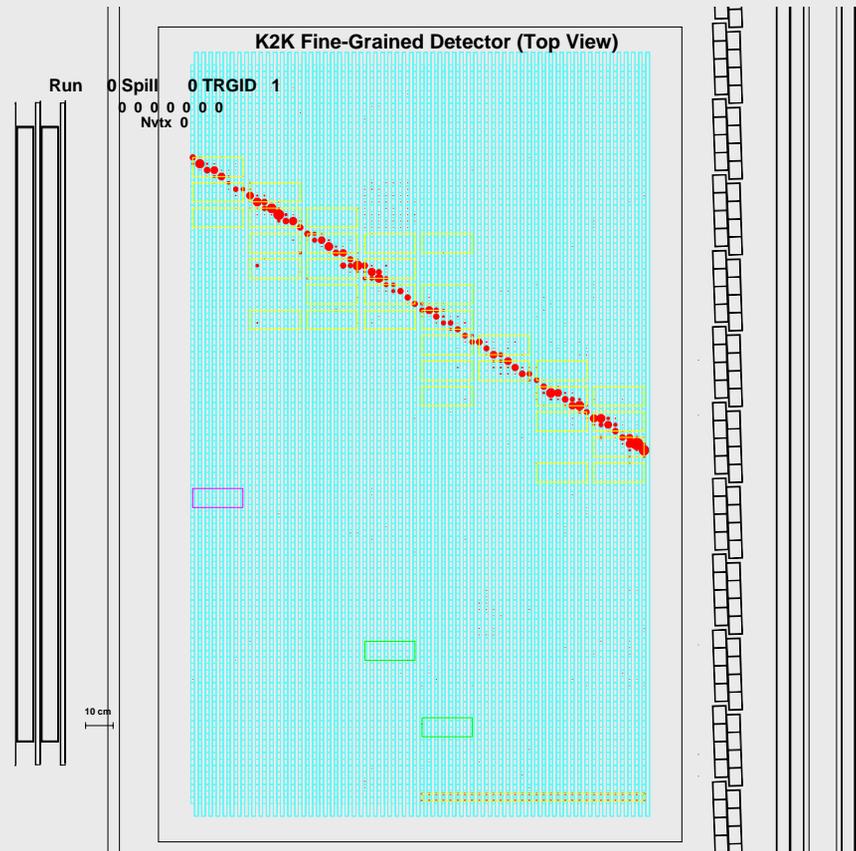
#channels :  $\sim 15,000$

# Cosmic Ray muon at SCIBAR

## SIDE View



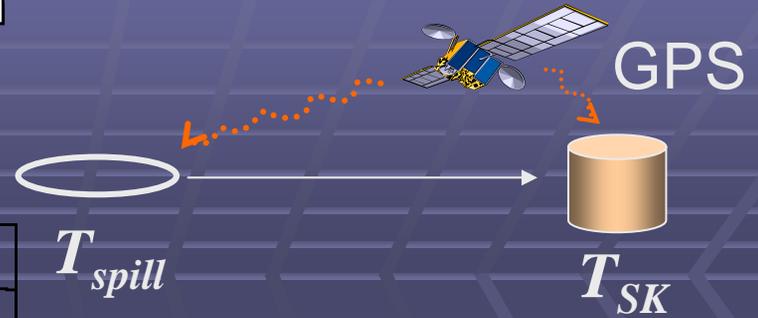
## TOP View



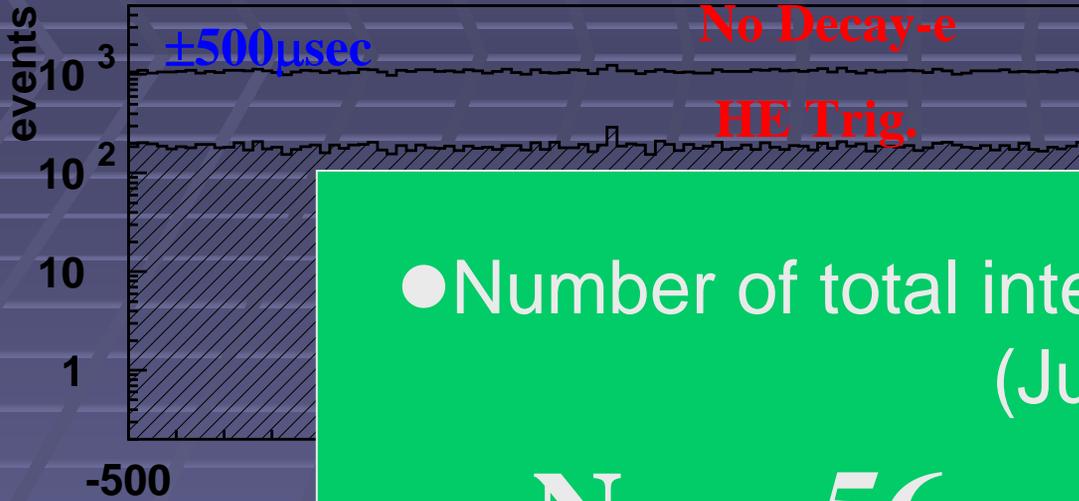
# Results of K2K-I

# Super-K Event selection

$$-0.2 \leq \Delta T \equiv T_{SK} - T_{Spill} - \text{TOF} \leq 1.3 \mu\text{sec}$$



$T_{Spill}$ : Abs. time of spill start  
 $T_{SK}$ : SK event  
 (SK to Kamioka)



● Number of total interactions  
 (Jun99-Jul01 )

$N_{obs} = 56$

$N_{exp} = 80.1^{+6.2}_{-5.4}$

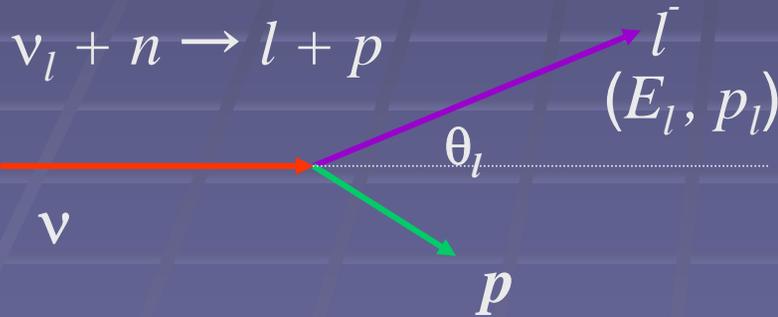
ed  
 (Water Detector)  
 ial Volume

atm.  $\nu$  BG  
 $<10^{-3}$  within  $1.5 \mu\text{s}$ .

# $E_\nu$ spectrum analysis

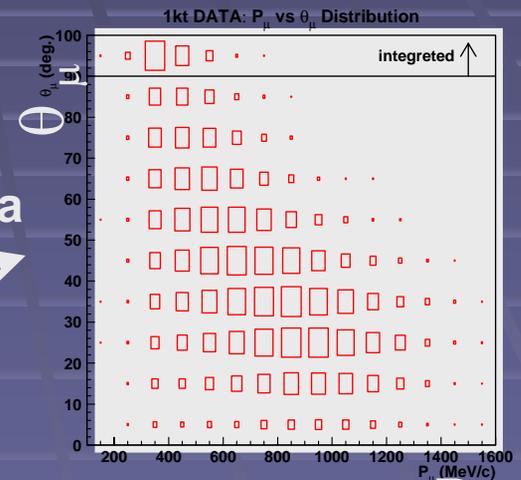
- Determination of expected  $\Phi(E_\nu)$  spectrum
  - Beam Monte Carlo
  - $\pi$ -monitor
  - Measurement of CC spectrum by near detector

CCQE interaction (2 body)



$$E_\nu = \frac{m_N E_l - m_l^2 / 2}{m_N - E_l + p_l \cos \theta_l}$$

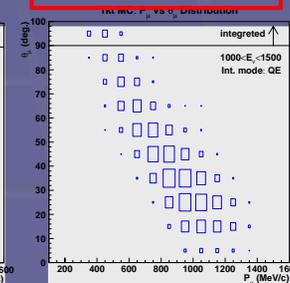
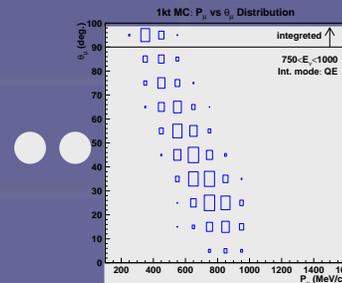
Spectrum fit  
MC Data



$E_\nu = 0.75 - 1.0$

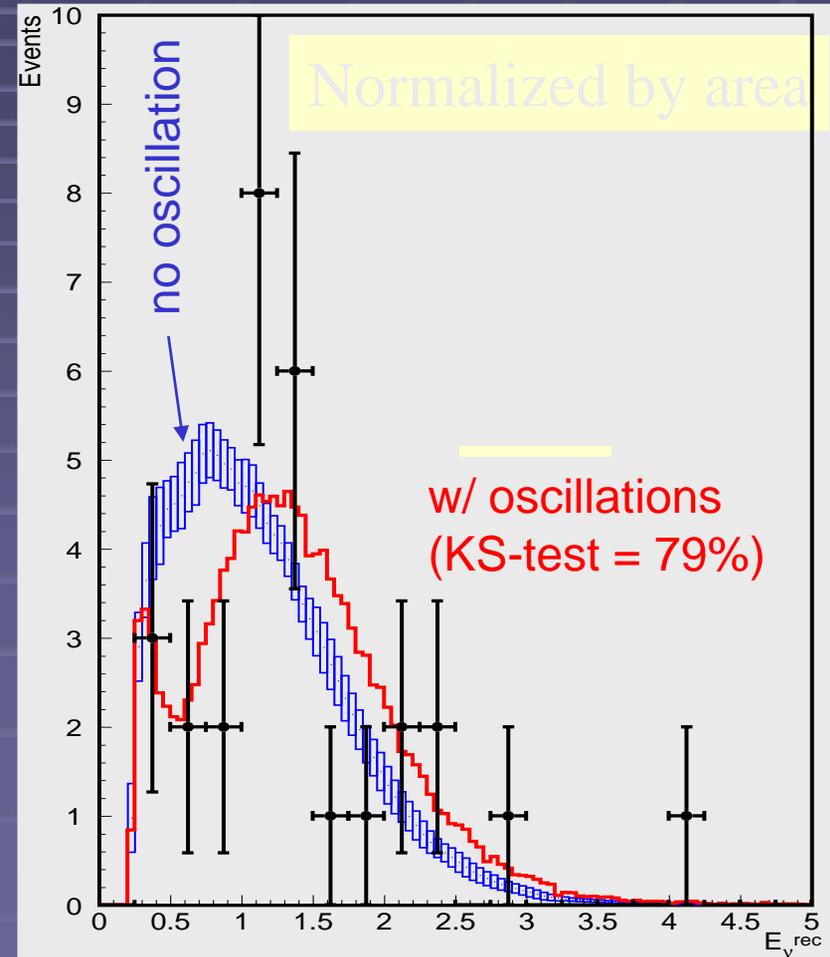
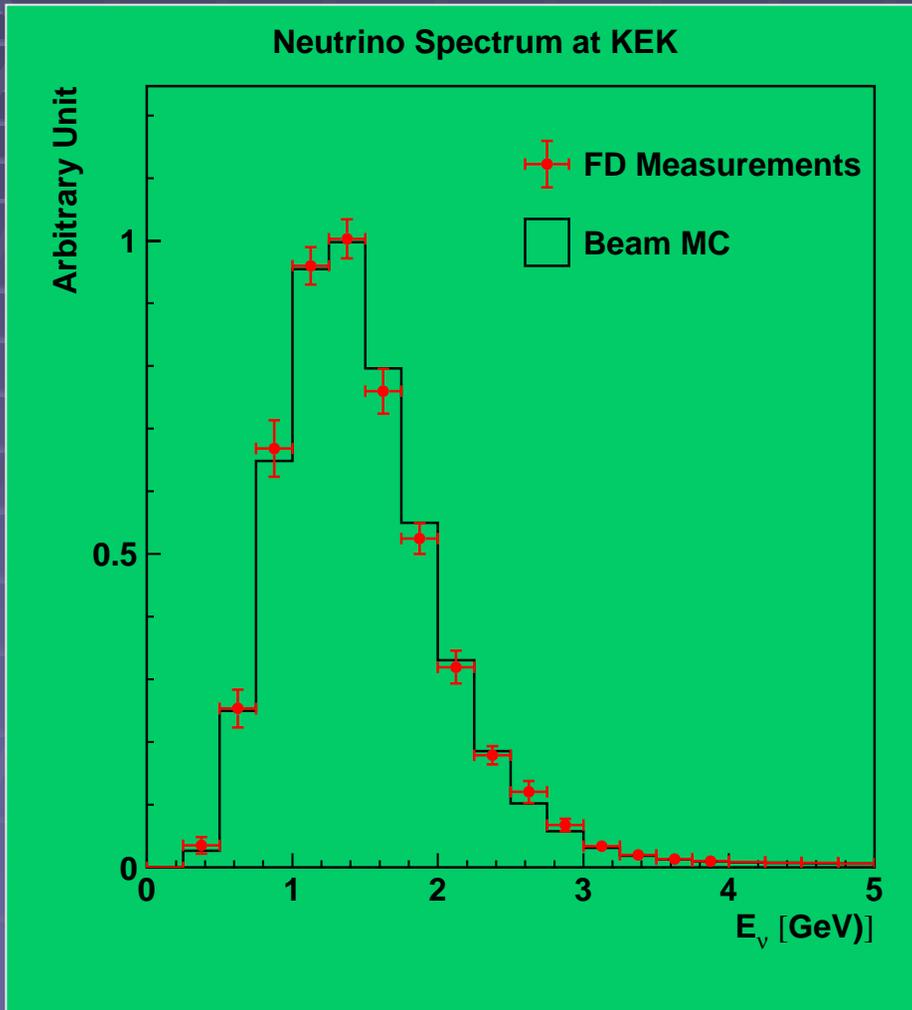
$1.0 - 1.5$

$P_\mu$



# $E_\nu$ spectrum results in K2K-I

- Reconstructed  $E_\nu$  shape of 1-RFC $\mu$  at SK (29 1-R events in Nov99-Jul01)

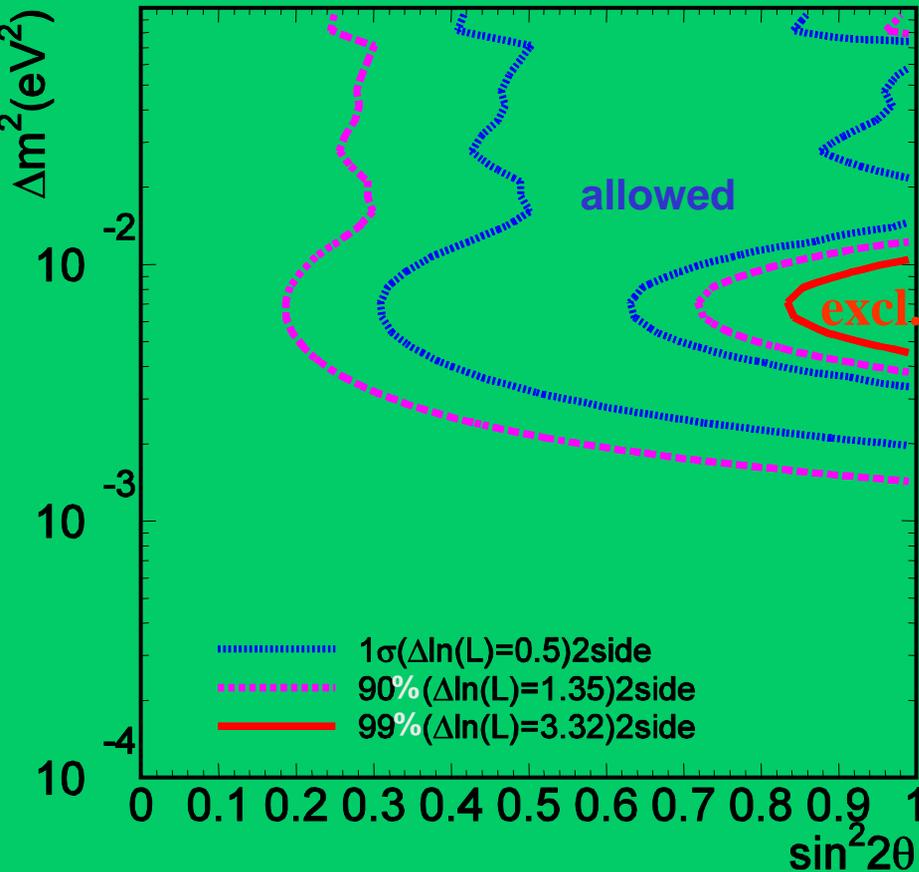


# Allowed regions and Null osc. probability

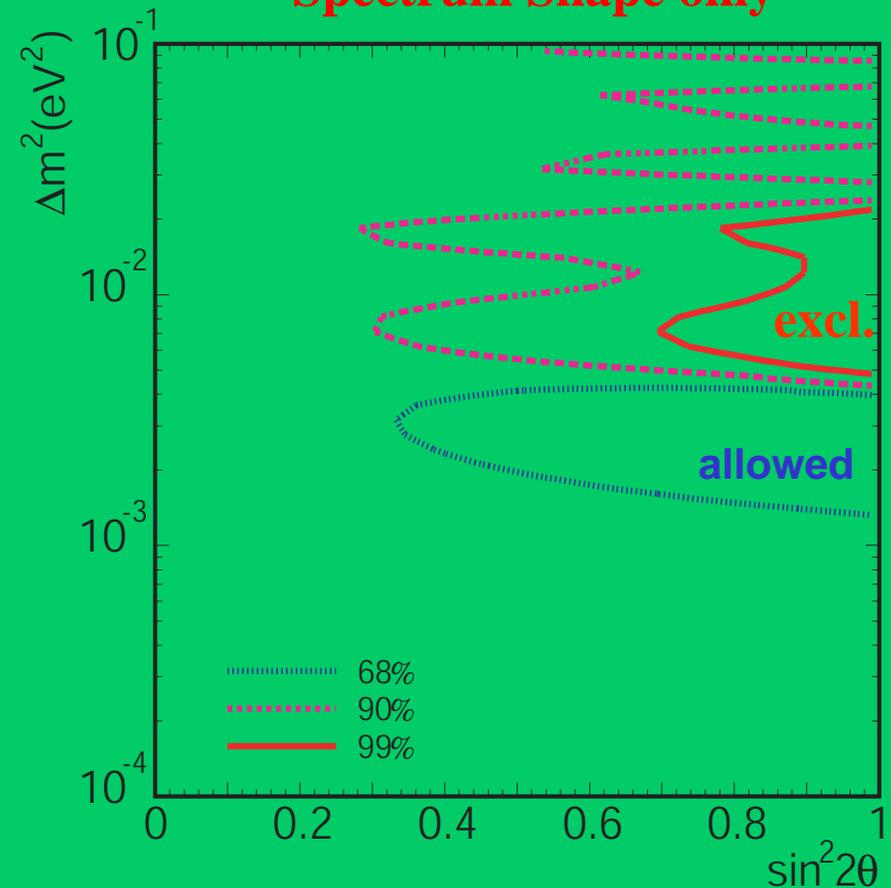
56FC events observed / 80.1 expected  
(Jun99-Jul01 data)

29 1-R FC $\mu$  events shape  
(Nov99-Jul01 data)

## Total no. of Events only

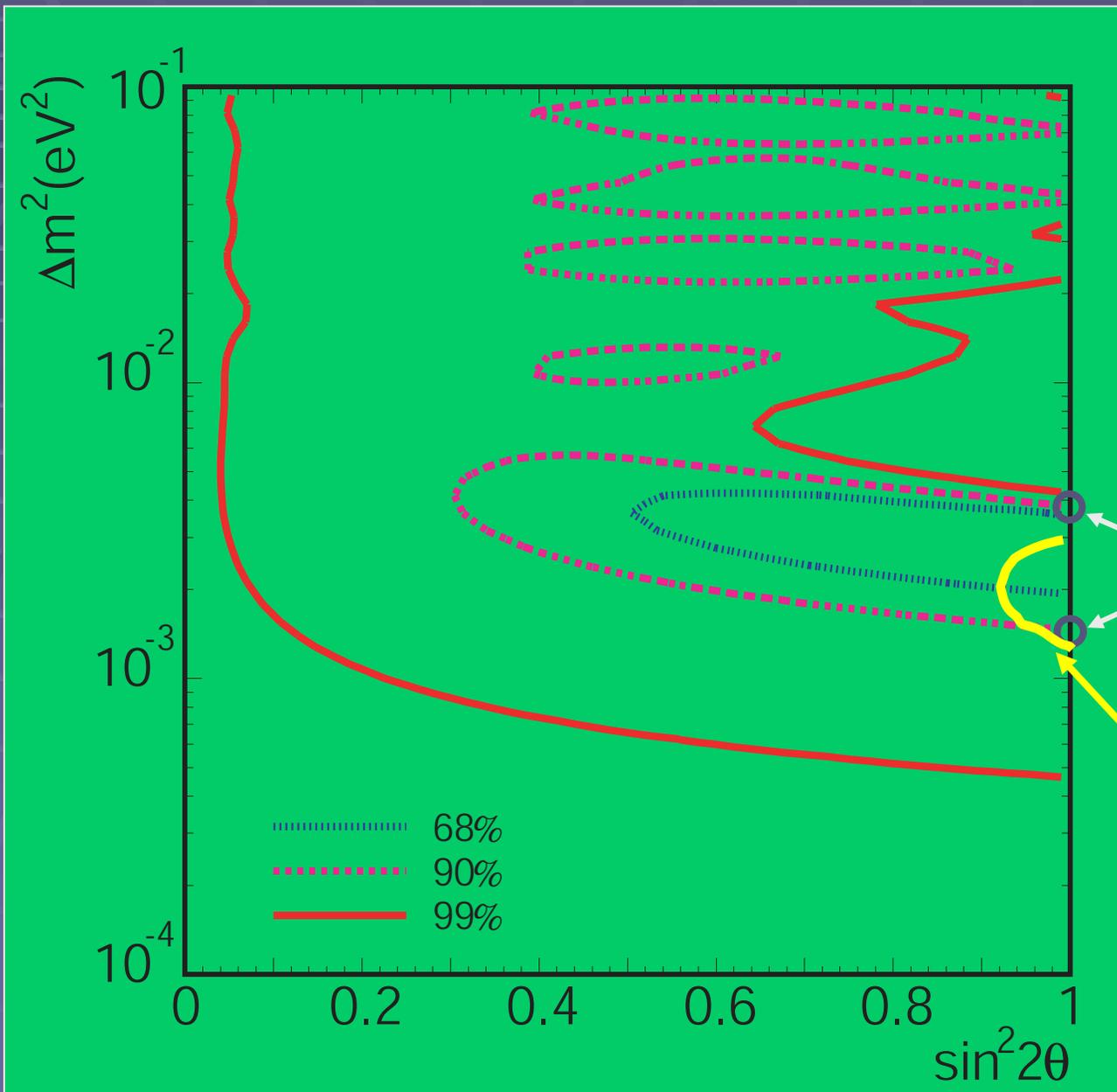


## Spectrum Shape only



**Shape and  $N_{SK}$  + Shape indicate consistent parameter region**

# Combined Allowed region (Shape+Norm) for K2K-I



● Null oscillation

⇒ < 1%

$\Delta m^2 =$   
 $1.5 \sim 3.9 \times 10^{-3} \text{eV}^2$   
@  $\sin^2 2\theta = 1$   
@ 90% CL

atm- $\nu$  results  
Consistent !



# Electron Candidate

★ Super Kamiokande ★

NUM	1
RUN	8071
SUBRUN	41
EVENT	5487540
DATE	99-Nov-6
TIME	15:17:5
TOT PE:	5647.
MAX PE:	39.2
NMHIT	2116
ANT-PE:	31.5
ANT-MX:	9.8
NMHITA:	29

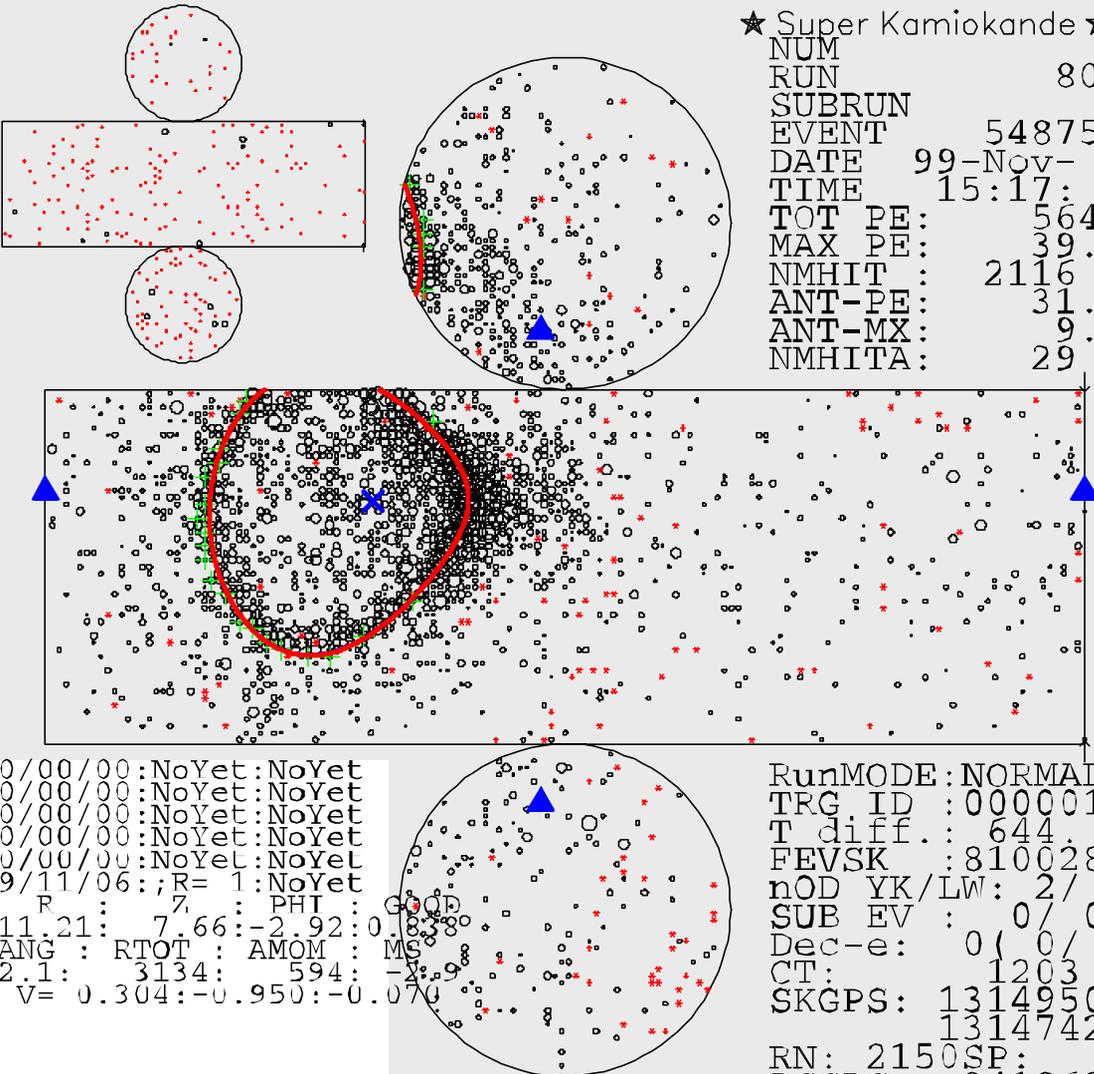
reconst. momentum

597 MeV/c

reconst. Ev

assuming  $\nu_e$  CCQE

612 MeV



```

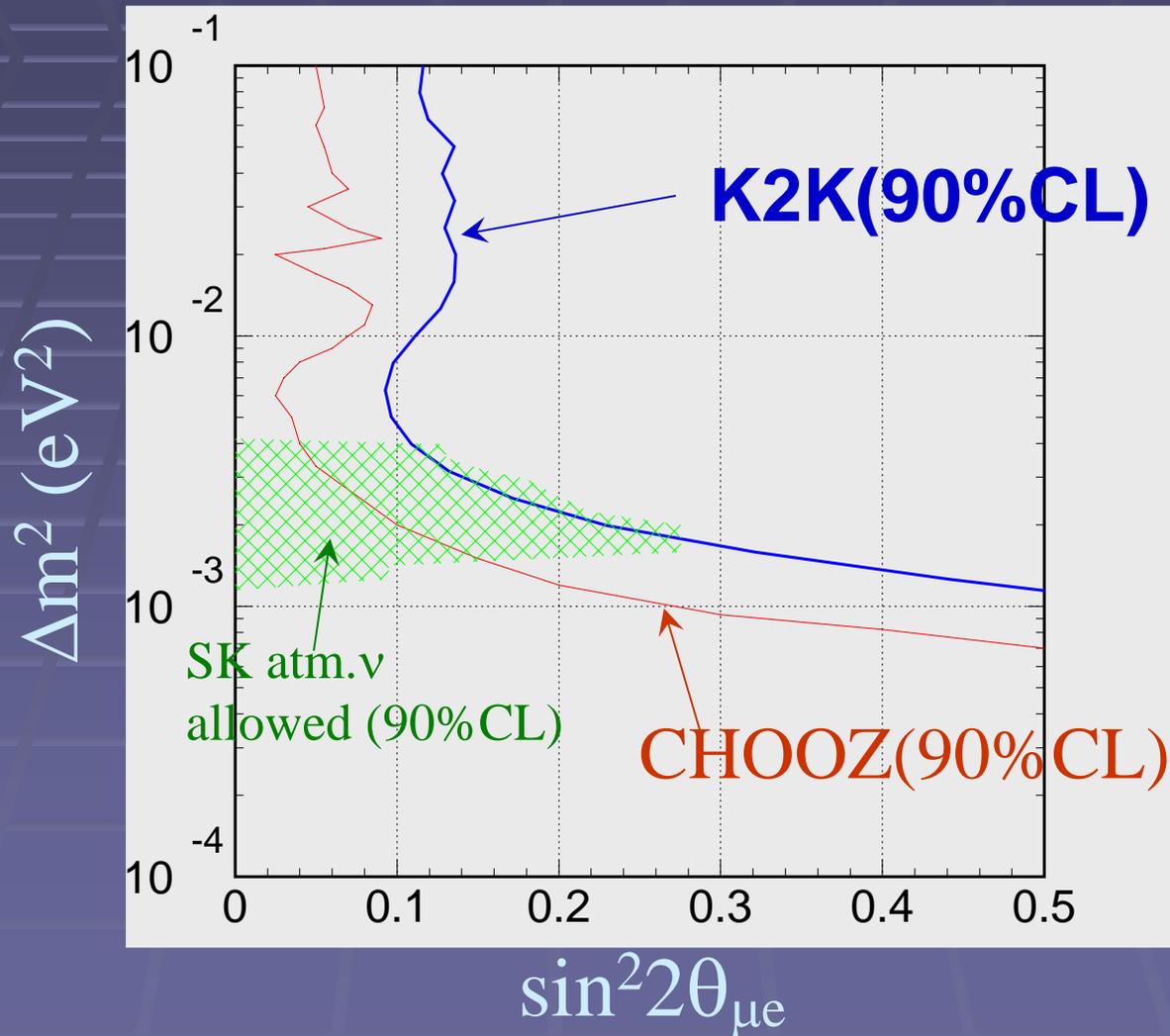
90/00/00:NoYet:NoYet
90/00/00:NoYet:NoYet
90/00/00:NoYet:NoYet
90/00/00:NoYet:NoYet
90/00/00:NoYet:NoYet
99/11/06:;R= 1:NoYet
  R      : 7      :PHI:
11.21: 7.66:-2.92:0.8380
CANG : RTOT : AMOM : MS
42.1: 3134: 594: -2.9
  V= 0.304:-0.950:-0.070
    
```

```

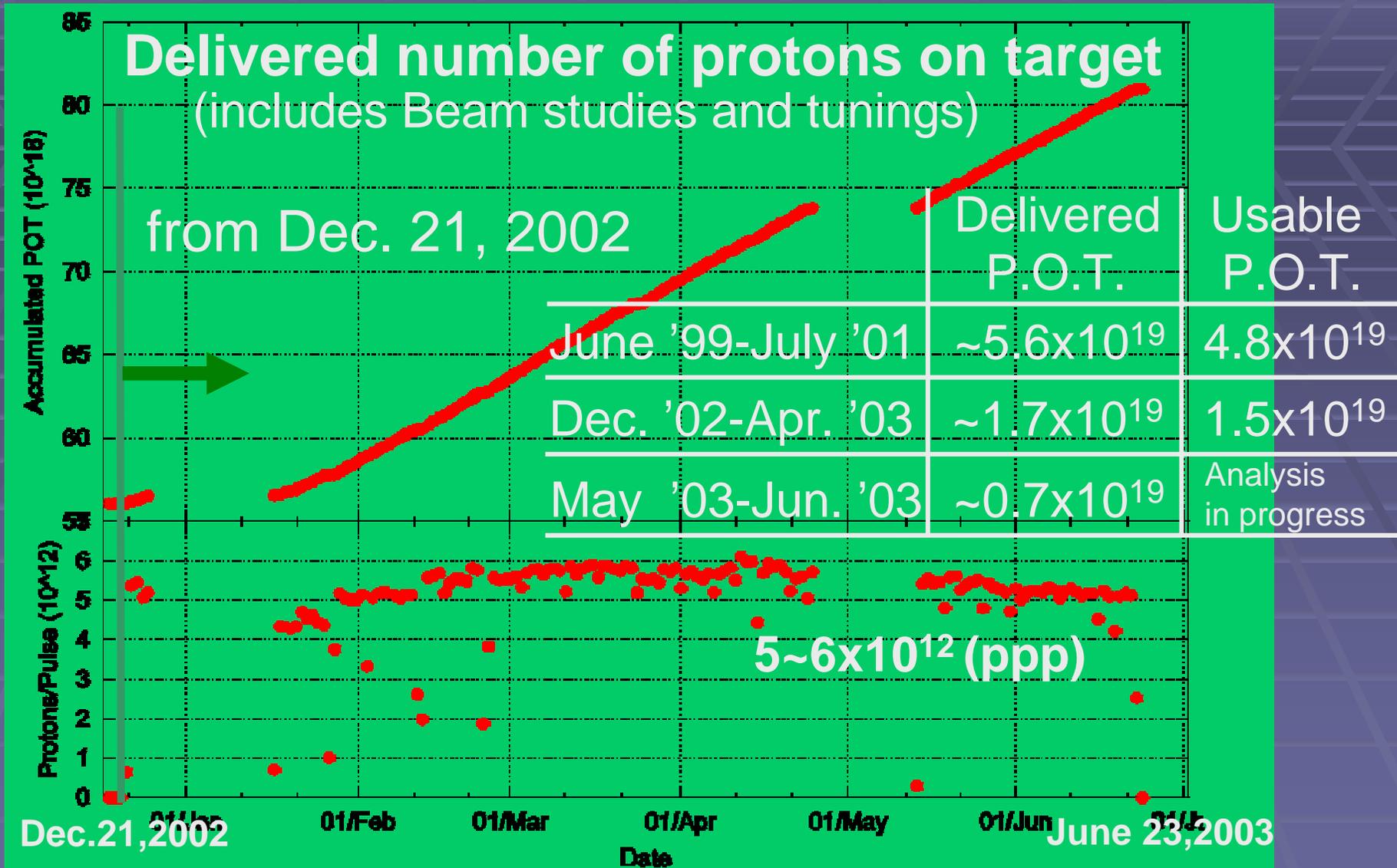
RunMODE: NORMAL
TRG ID : 00000111
T diff.: 644.
FEVSK  : 81002803
nOD YK/LW: 2/ 3
SUB EV  : 0/ 0
Dec-e: 0( 0/ 0/
CT: 1203
SKGPS: 131495094
      131474205
RN: 2150SP:
PSGPS: 94186902
      92767476
GPSDIF: 0.41
    
```

Comnt;

# Allowed region for $\sin^2 2\theta_{\mu e}$

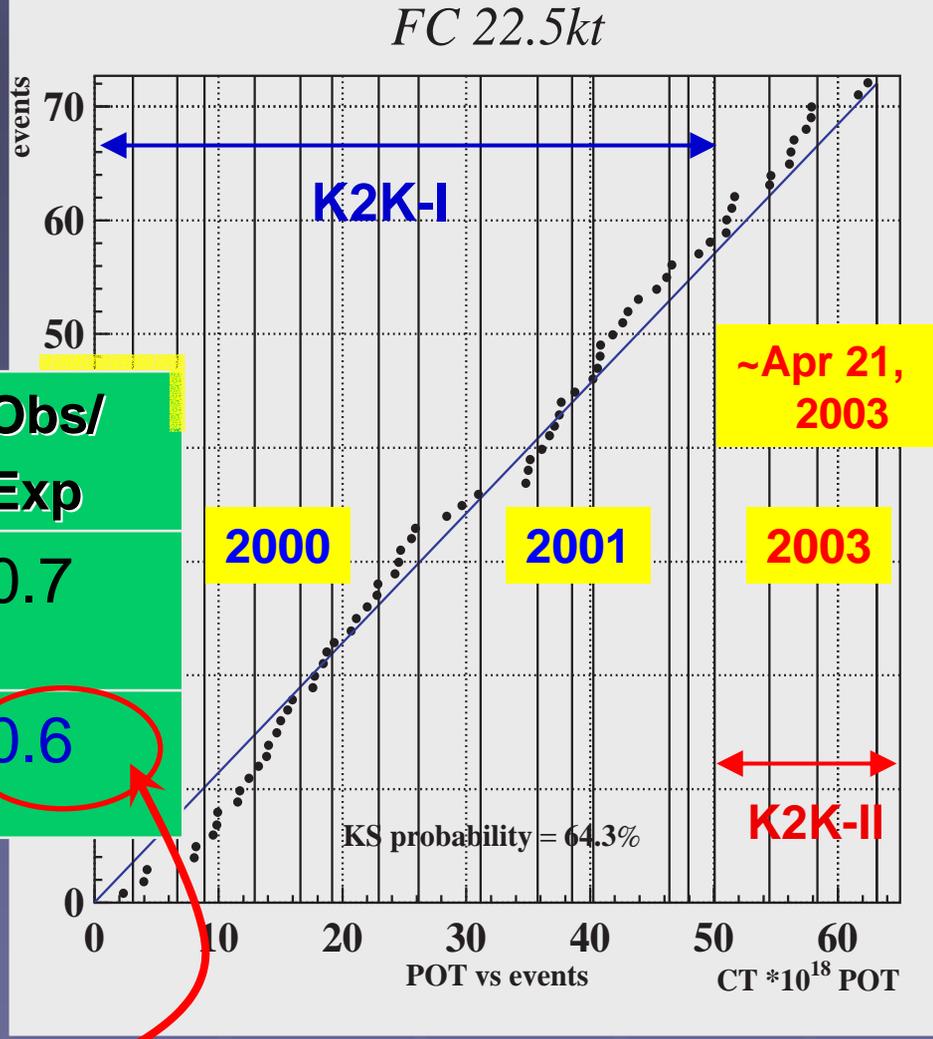
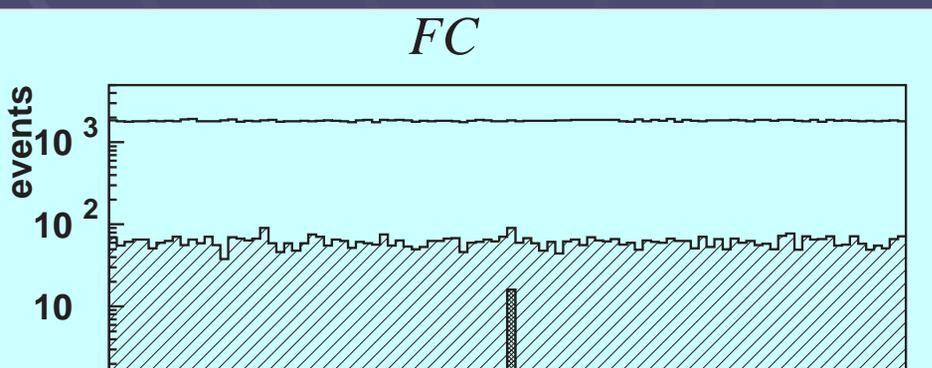


# Status of K2K-II



Very stable and no serious problems

# Status of K2K-II K2K-II Preliminary



	Obs.	Exp.	Obs/ Exp
1999-2001 (K2K-I)	56	80.1	0.7
2003 Jan ~ Apr (K2K-II)	16	26	0.6



**K2K-II experiment observed consistent reduction rate**

# Summary

- Atmospheric  $\nu$  results from SK-I
  - Finalization for SK-I data is going on
  - $\Delta m^2 = 1.3 \sim 3.0 \times 10^{-3} \text{eV}^2$ ,  $\sin^2 2\theta > 0.92$  @ 90%CL
- K2K results
  - K2K-I results (Total events +  $E_\nu$  spectrum )
  - Null oscillation probability is less than 1%
  - $\Delta m^2 = 1.5 \sim 3.9 \times 10^{-3} \text{eV}^2$  for  $\sin^2 2\theta = 1$  @ 90%CL
- Sk-II / K2K-II successfully resumed
  - K2K-II observe consistent  $\nu$  rate with K2K-I