

# **NEW PERSPECTIVE ON GALAXY CLUSTERING AND COSMOLOGY: GENERAL RELATIVISTIC EFFECTS**

**JAIYUL YOO**

**INSTITUTE for THEORETICAL PHYSICS, UNIVERSITY of ZÜRICH  
LAWRENCE BERKELEY LABORATORY, UNIVERSITY of CALIFORNIA, BERKELEY**

**Korea Institute for Advanced Study, June, 28, 2011**

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# **I. INTRODUCTION:**

*Galaxies as a Cosmological Probe – What is the Problem?*

# Dark Energy Surveys

- **current and future surveys:**
  - **Baryonic Oscillation Spectroscopic Survey**
  - **Dark Energy Survey**
  - **Panoramic Survey Telescope & Rapid Response System**
  - **Hobby-Eberly Telescope Dark Energy Experiment**
  - **Wide Field Multi-Object Spectroscopy**
  - **Large Synoptic Survey Telescope**
- **future space missions:**
  - **EUCLID, Supernova Acceleration Probe**
  - **Cosmic Inflation Probe**
  - **Wide-Field Infrared Survey Telescope**

## Legend of Galaxies

- **current and future dark energy surveys:**
  - *better precision and larger scales!*
- **galaxies as cosmological probes:**
  - BAO signature:  $D_A$ ,  $\Omega_b/\Omega_m$ ,  $k \sim 0.06$  h/Mpc
  - galaxy power spectrum:  $n_s$ ,  $\Omega_k$ ,  $\Omega_m h$ ,  $k \sim 0.01$  h/Mpc
  - primordial non-Gaussianity:  $f_{NL}$ ,  $n_s$ ,  $k \sim 0.001$  h/Mpc

*a “clean” cosmological probe in linear regime*

*BUT is our faith well founded?*

# Cosmological Probe

- *precision cosmology!*
- **galaxies trace underlying matter**
  - **biased tracer:**  $\delta_g = b \delta_m$
  - **z-space distortion:**  $\delta_g = b \delta_m - \frac{1+z}{H} \frac{\partial V}{\partial r}$
  - **gravitational lensing:**  $\delta_g = b \delta_m + (5p - 2) \kappa$
  - **contributions are added in *adhoc* manner!**

*is this everything? or are there more contributions?*  
*we need unified treatments!*

## Relativistic Perspective

- **theoretical inconsistency in galaxy clustering**

- **standard description:**  $\delta_g = b \delta_m - \frac{1+z}{H} \frac{\partial V}{\partial r}$

- **synchronous gauge (e.g., CMBFAST, CAMB)**

- **free falling frame**  $\psi = V = 0$

- **Poisson equation**  $\nabla^2 \psi \sim \delta_m \sim 0 \quad ?$

*galaxy clustering is based on Newtonian description!*

## Gauge Issues

- **theoretical predictions in cosmology**
    - compute perturbations such as  $\delta_m$ ,  $\psi$ ,  $P_m(k)$ ,  $\dots$
    - compare to observable quantities
    - perturbations are ***gauge-dependent***
    - ***so are many theoretical descriptions!***
  - **observable quantities**
    - gauge-invariance is a necessary condition
- theoretical predictions vs observables: gauge issue!***

## **II. LARGE-SCALE STRUCTURE:**

*General Relativistic Description – New Perspective*

# Cosmology

- modern cosmology: general theory of relativity
- **cosmological framework:**
  - described by Einstein equations
  - homogeneous & isotropic FLRW universe
  - inflation, big bang nucleosynthesis, CMB
- *galaxies in cosmological framework?*
  - Newtonian, no GR description
  - automatic disqualification?

# Galaxies in General Relativity

- *what are observables?*
- geodesic equations of photons from galaxies
- **time component:** *Sachs-Wolfe* effect

- observed redshift

$$1 + z_{\text{obs}} = (1 + z) \left[ 1 + V(z) - V(0) - \psi(z) + \psi(0) - \int_0^r dr' (\dot{\psi} - \dot{\phi}) \right] .$$

- **spatial component:** *gravitational lensing* effect
  - observed position  $\hat{n} = (\theta, \phi)$
  - lensing displacement  $(\delta r, \delta \theta, \delta \phi)$
  - magnification  $\mu \simeq 1 + 2\kappa$

# Effects on Galaxies

- **construct a galaxy fluctuation field:**

- **total number of observed galaxies**  $N_{\text{tot}}$
- **observed volume**  $dV_{\text{obs}}$  **given**  $(z_{\text{obs}}, \hat{n})$
- **fluctuation field**  $\delta_{\text{obs}} = \frac{n_{\text{obs}}}{\langle n_{\text{obs}} \rangle} - 1$

- **relation to true number density:**

- **number conservation**  $N_{\text{tot}} = n_{\text{true}} dV_{\text{true}} = n_{\text{obs}} dV_{\text{obs}}$
- **volume element**  $dV_{\text{obs}} = \frac{r^2(z_{\text{obs}})}{H(z_{\text{obs}})} dz_{\text{obs}} d\Omega_{\text{obs}}$
- **note**  $z_{\text{true}} \neq z_{\text{obs}}$  ,  $d\Omega_{\text{true}} \neq d\Omega_{\text{obs}}$

# Unified Treatment

- **observable:**  $N_{\text{tot}} = n_{\text{true}} dV_{\text{true}} = n_{\text{obs}} dV_{\text{obs}}$
- **volume effects:**
  - **redshift-space distortion:**  $\frac{\partial z}{\partial z_{\text{obs}}} \simeq \frac{1+z}{H} \frac{\partial V}{\partial r}$
  - **lensing magnification:**  $\frac{\partial \Omega}{\partial \Omega_{\text{obs}}} \frac{\partial f}{\partial f_{\text{obs}}} \simeq \frac{1}{\mu^2} = 1 - 4 \kappa$
- **source effects:**
  - **magnification bias:**  $\bar{n}_{\text{obs}}(f_{\text{obs}}) \simeq \bar{n}(f_{\text{obs}}/\mu)$
- *complete description of different effects*

# Subtle Issues

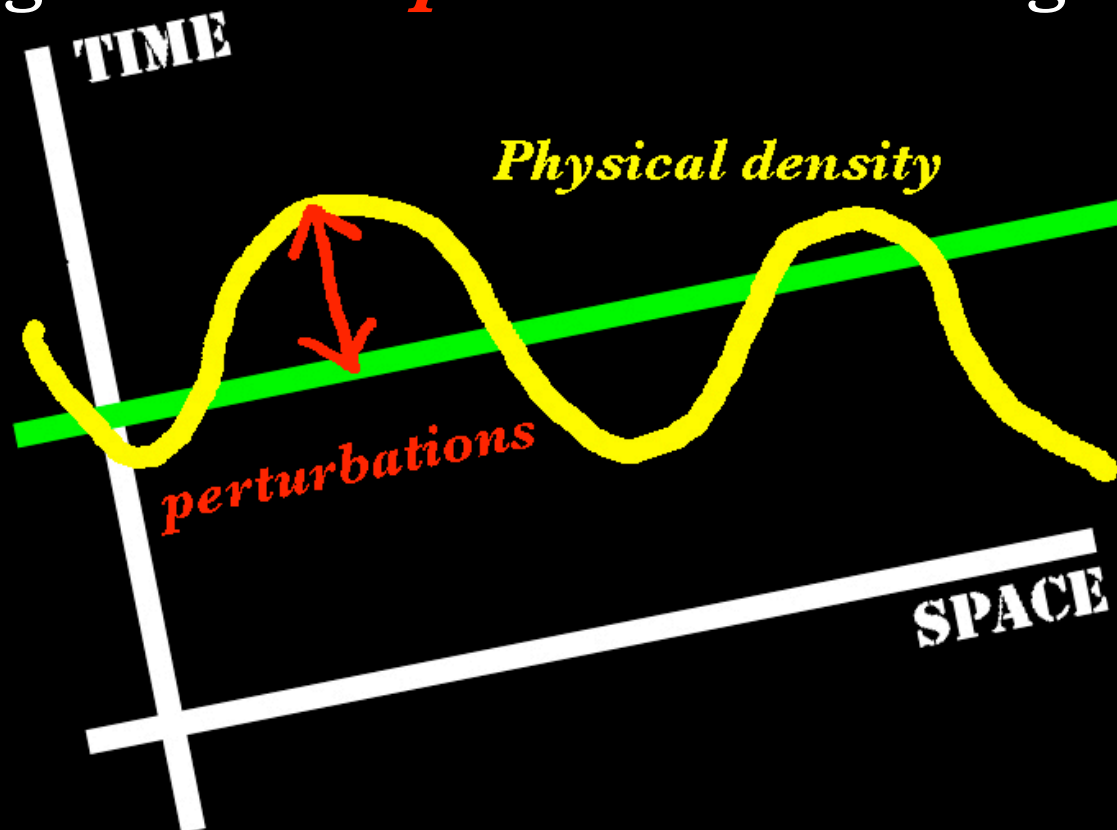
- *what are “true” volume and number density?*
- **true number density:**  $n_{\text{true}} = n_{\text{phy}}$ 
  - number density in local inertial frame
- **true volume:**  $dV_{\text{true}} = \frac{r^2(z_{\text{true}})}{H(z_{\text{true}})} dz_{\text{true}} d\Omega_{\text{true}} ?$ 
  - $z_{\text{true}}$  ,  $\hat{s}$  in FLRW universe *not really “true!”*
  - at what coordinate system (*gauge*)?  
e.g.,  $z_{\text{true}} = z(t)$  ,  $\bar{\rho}_m = \bar{\rho}_m(t)$
- **fully relativistic theory:**
  - quest for true volume
  - better understanding of gauge issues

# Correspondence

- **cosmological perturbations:**
  - inhomogeneous physical spacetime
  - homogeneous *fictitious* background

# Gauge Freedom

- **general covariance in GR:**
  - free to choose a coordinate system
  - change in *correspondence* to background



## Gauge Issues

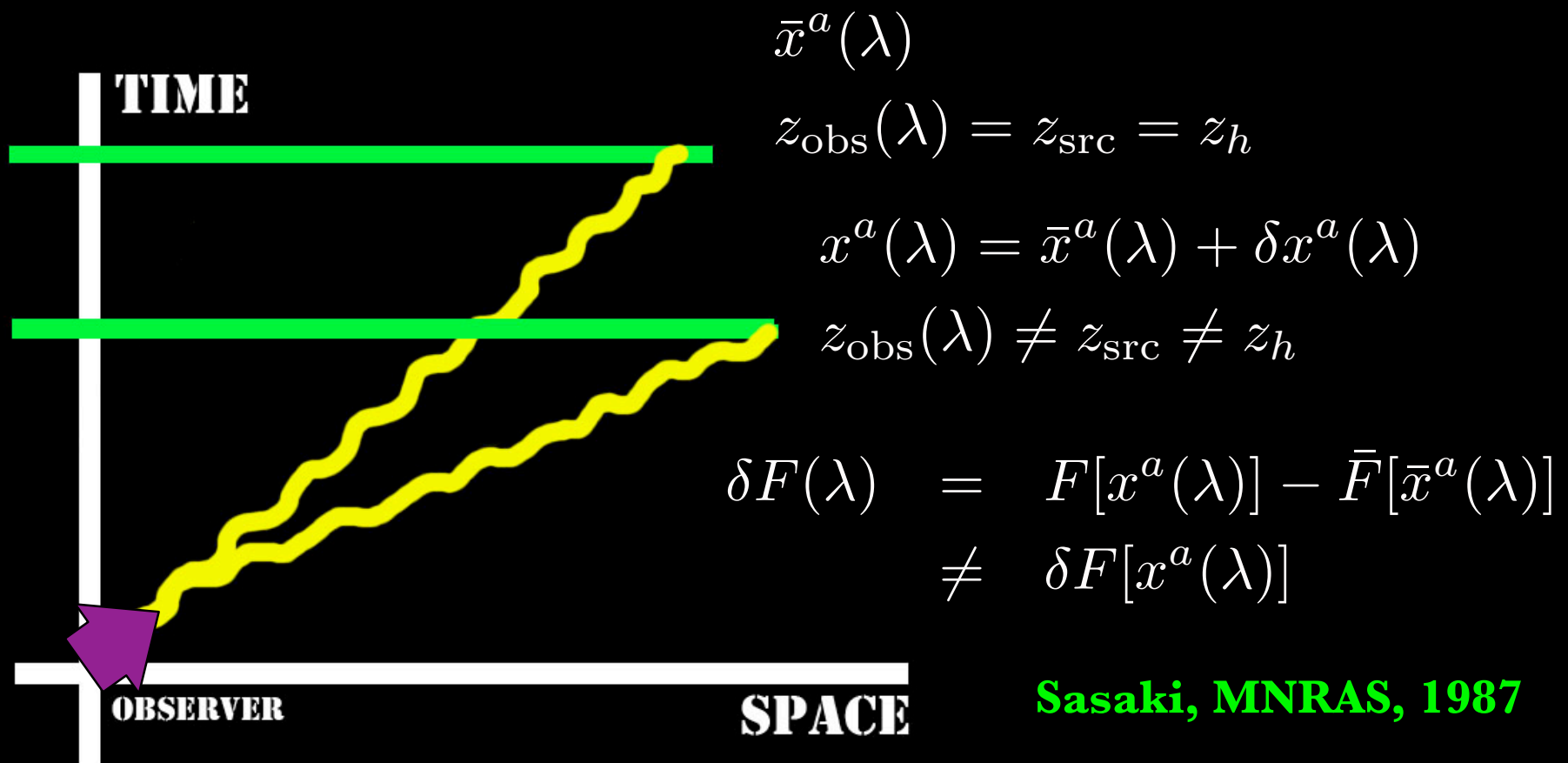
- **gauge-dependent quantities**
  - change its value depending on coordinates
  - *cannot be directly associated with observables!*
- **theoretical descriptions of observables**
  - should be gauge-invariant
  - *but surprisingly not for most!*
- **Newtonian limit**
  - unambiguous hypersurface of simultaneity

## More Subtle Issues

- **observed mean** vs **ensemble average**
  - observed mean is *gauge-invariant*
  - ensemble average is *gauge-dependent*
- spatial gauge transformation (*unphysical*)
  - mapping is invariant, time is fixed
- **observer position**
  - also gauge-dependent

# Complications

- **perturbations along photon geodesic:**
  - **observed angle and redshift**



**Sasaki, MNRAS, 1987**

# Observed Redshift

- **observed redshift:**

$$1 + z_{\text{obs}} = (1 + z) \left[ 1 + V(z) - V(0) - \psi(z) + \psi(0) - \int_0^r dr' (\dot{\psi} - \dot{\phi}) \right].$$

- **observer's point of view:**

- true redshift  $z$  from observed redshift  $z_{\text{obs}}$

- **new perspective:**

- true redshift is *fictitious & gauge-dependent*
- in another coordinate  $z$ ,  $V$ ,  $\psi$  *change!*

# Observed Redshift

- **observed redshift:**

$$1 + z_{\text{obs}} = (1 + z) \left[ 1 + V(z) - V(0) - \psi(z) + \psi(0) - \int_0^r dr' (\dot{\psi} - \dot{\phi}) \right] .$$

$$= (1 + z^s) \left[ 1 + \frac{1}{6} \int_0^r dr' (\dot{h} + 3 \dot{h}_{\alpha\beta}^{\parallel} e^{\alpha} e^{\beta}) \right] \quad \textbf{(synchronous)}$$

- **gauge transformation:**

- **true redshift**  $z \neq z^s$
- **true volume element**  $\frac{r^2(z)}{H(z)} \neq \frac{r^2(z^s)}{H(z^s)}$
- **observed redshift is *gauge-invariant***  $z_{\text{obs}} = z_{\text{obs}}^s$

# Gravitational Lensing

- **observed position:**  $\hat{n} = (\theta, \phi)$
- **lensing displacements:**  $(\delta r, \delta\theta, \delta\phi)$  &  $\delta\tau$
- **true position:**  $\hat{s} = (\theta + \delta\theta, \phi + \delta\phi)$
- **magnification:**  $\mu \simeq \left| \frac{d^2 \hat{n}}{d^2 \hat{s}} \right| = 1 + 2 \kappa$
- **gauge-dependent quantities:**  $\delta r, \delta\theta, \delta\phi, \hat{s}, \kappa$ 
  - *standard weak lensing is gauge-dependent!*

# Magnification

- definition is *inaccurate!*
- **coordinate-independent definition:**
  - luminosity in local inertial frame  $L$
  - measure flux & redshift  $f_{\text{obs}}$  &  $z_{\text{obs}}$
  - magnification is *physical!*

$$\mu = f_{\text{obs}} \left( \frac{L}{4\pi D_L^2(z_{\text{obs}})} \right)^{-1}$$

- in a homogeneous universe using  $z_{\text{obs}}$ , not  $z_{\text{true}}$

$$D_L(z_{\text{obs}}) = (1 + z_{\text{obs}}) r(z_{\text{obs}})$$

# Luminosity Distance

- **observed flux and intrinsic luminosity:**
  - **includes Sachs-Wolfe and lensing effects**

$$\mathcal{D}_L(z_{\text{obs}}) = \sqrt{\frac{L}{4\pi f_{\text{obs}}}} \equiv D_L(z_{\text{obs}})(1 + \delta\mathcal{D}_L) \quad \text{Sasaki, MNRAS, 1987}$$
$$= D_L(z_{\text{obs}}) \left[ 1 + \frac{\delta\lambda}{r_s} - \frac{1}{2} \int_0^{r_s} dr \delta\theta \right]$$

- **shift in affine parameter**  $\delta\lambda \sim \delta z$
- **distortion in wave vector expansion**  $\delta\theta \sim \kappa$
- **gauge-invariant**  $\delta\mathcal{D}_L$

# Luminosity Distance

- **relation to magnification:**

$$\mu = f_{\text{obs}} \left( \frac{L}{4\pi D_L^2(z_{\text{obs}})} \right)^{-1} = \left( \frac{D_L(z_{\text{obs}})}{\mathcal{D}_L(z_{\text{obs}})} \right)^2$$
$$= 1 - 2 \delta\mathcal{D}_L \simeq 1 + 2 \kappa$$

- **prevalent gauge issues in cosmology:**

- observed magnification is *physical*
- *but* usual parametrization is *gauge-dependent*
- **source effect**  $\bar{n}_{\text{obs}} = \bar{n}[f_{\text{obs}}(1 + 2\delta\mathcal{D}_L)]$

# Observed Number of Galaxies

- *we still need “true” volume!*
- **total number of observed galaxies:**

- **observables**  $N_{\text{tot}}, \hat{n} = (\theta, \phi), z_{\text{obs}}$

$$\begin{aligned} N_{\text{tot}} &= \int dz_{\text{obs}} d\Omega_{\text{obs}} n_{\text{obs}} \frac{r^2(z_{\text{obs}})}{(1 + z_{\text{obs}})^3 H(z_{\text{obs}})} \\ &= \int n_{\text{phy}} dV_{\text{phy}} \end{aligned}$$

- **physical volume:**  $dV_{\text{phy}}$ 
  - **occupied by observed galaxies**
  - *trace backward photon geodesic!*

# Matias's Magic

- **integral of 3-form in 4D spacetime manifold:**
  - **observables**  $z_{\text{obs}}, \theta_{\text{obs}}, \phi_{\text{obs}}$
  - **photon geodesic path**  $x^a(\lambda) = \bar{x}^a(\lambda) + \delta x^a(\lambda)$
  - **Sachs-Wolfe and gravitational lensing effects**
  - **distortion in local Lorentz frame**
  - ***manifestly gauge-invariant***

$$N_{\text{tot}} = \int \sqrt{-g} \, n_{\text{phy}} \, \varepsilon_{abcd} \, u^d \, \frac{\partial x^a}{\partial z_{\text{obs}}} \frac{\partial x^b}{\partial \theta_{\text{obs}}} \frac{\partial x^c}{\partial \phi_{\text{obs}}} \, dz_{\text{obs}} \, d\theta_{\text{obs}} \, d\phi_{\text{obs}}$$

**Levi-Civita symbol**  $\varepsilon_{abcd}$  , **comoving velocity**  $u^a$

# Observed Number of Galaxies

- **fun** and/or **pain** in **perturbation expansion!**

$$\begin{aligned}
 N_{\text{tot}} &= \int \sqrt{-g} \, n_{\text{phy}} \, \varepsilon_{abcd} \, u^d \, \frac{\partial x^a}{\partial z} \frac{\partial x^b}{\partial \theta} \frac{\partial x^c}{\partial \phi} \, dz \, d\theta \, d\phi \\
 &= \int n_{\text{phy}} \frac{r^2 \sin \theta}{(1+z)^3 H} \, dz \, d\theta \, d\phi \left[ 1 + 3D + V + 2 \frac{\delta r}{r} \right. \\
 &\quad \left. + H \frac{\partial}{\partial z} \delta r + \left( \cot \theta + \frac{\partial}{\partial \theta} \right) \delta \theta + \frac{\partial}{\partial \phi} \delta \phi + \frac{\bar{r}^2}{r^2} H \frac{\partial \bar{r}}{\partial z} \right] \\
 &\equiv \int n_{\text{obs}} \frac{r^2 \sin \theta}{(1+z)^3 H} \, dz \, d\theta \, d\phi
 \end{aligned}$$

- subscript “**obs**” is omitted!

# Observed Number Density

- **so far, we have**

- **volume effects:**  $n_{\text{obs}} = n_{\text{phy}} \left( 1 + \sum_{\mu} \delta_{\mu} \right)$

- **source effects:**  $\bar{n}_{\text{obs}} \rightarrow \bar{n}_{\text{phy}} [f_{\text{obs}}(1 + 2\delta\mathcal{D}_L)]$

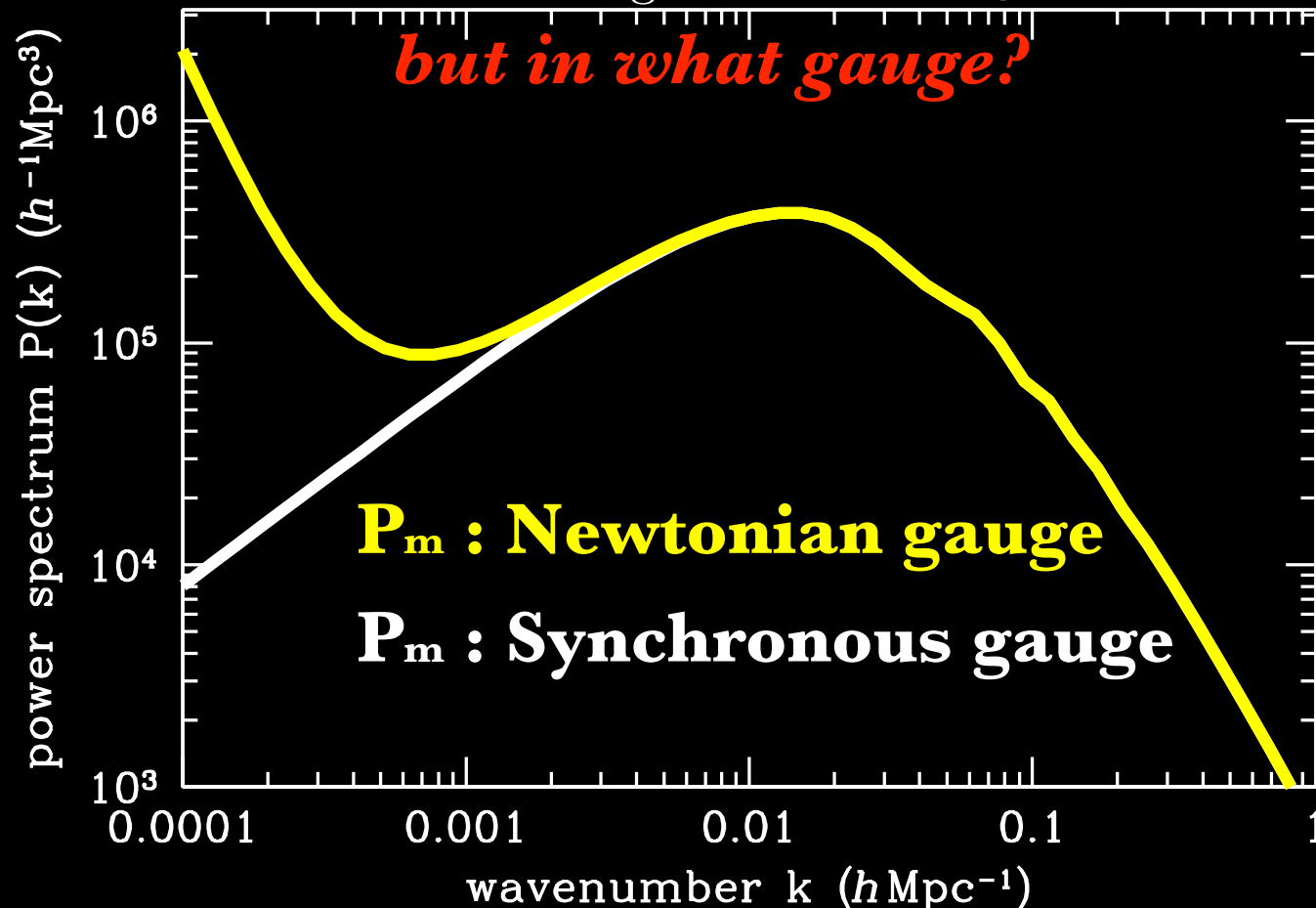
$$dn_{\text{phy}}/dL \propto L^{-s} \qquad = \bar{n}_{\text{phy}}(L_{\text{thr}})(1 - 5p \delta\mathcal{D}_L)$$

$$p = 0.4 (s - 1)$$

**BUT** *why do we care about **galaxies?***

# Galaxy Bias

- galaxies trace underlying matter!
- standard relation  $\delta_{\text{gal}} = b \delta_m$  ( $\delta_m$  : perturbation)



# Galaxy Bias

- **galaxy formation**
  - described in a local coordinate (proper time)

$$n_{\text{phy}} = F[\rho_m] \quad \rightarrow \quad n_{\text{phy}} = F[\rho_m, t_p] ,$$

- **time slicing in observation**  $1 + z_{\text{obs}} = (1 + z)(1 + \delta z)$ 
  - *observed redshift* defines simultaneity
  - measured at observed redshift

$$n_{\text{phy}} = \bar{n}_{\text{phy}}(z_{\text{obs}}) [1 + b \delta_m^{(v)} - e \delta z^{(v)}]$$

$$e = \frac{d \ln \bar{n}_{\text{phy}}}{d \ln(1 + z)} , \quad 1 + z_{\text{obs}} = (1 + z)(1 + \delta z)$$

# Cosmological Probe

- accurate relation to underlying matter*

- most cases :**  $\delta_g = b \delta_m$

- prudent work :**  $\delta_g = b \delta_m - \frac{1+z}{H} \frac{\partial V}{\partial r}$

- best efforts so far :**  $\delta_g = b \delta_m - \frac{1+z}{H} \frac{\partial V}{\partial r} + (5p - 2) \kappa$

- this work :** 
$$\delta_g = b \delta_m^{(v)} - e \delta z^{(v)} + \alpha_\chi + 2 \varphi_\chi + V - C_{\alpha\beta} e^\alpha e^\beta$$

$$+ 3 \delta z_\chi + 2 \frac{\delta \mathcal{R}}{r} - H \frac{\partial}{\partial z} \left( \frac{\delta z_\chi}{\mathcal{H}} \right) - 5p \delta \mathcal{D}_L - 2 \mathcal{K} ,$$

Yoo, Fitzpatrick, Zaldarriaga, PRD, 2009

Yoo, PRD, 2010

# Scales of Interest

- **Synchronous**  
(CMBFast, CAMB)

$$P_\delta(k)$$

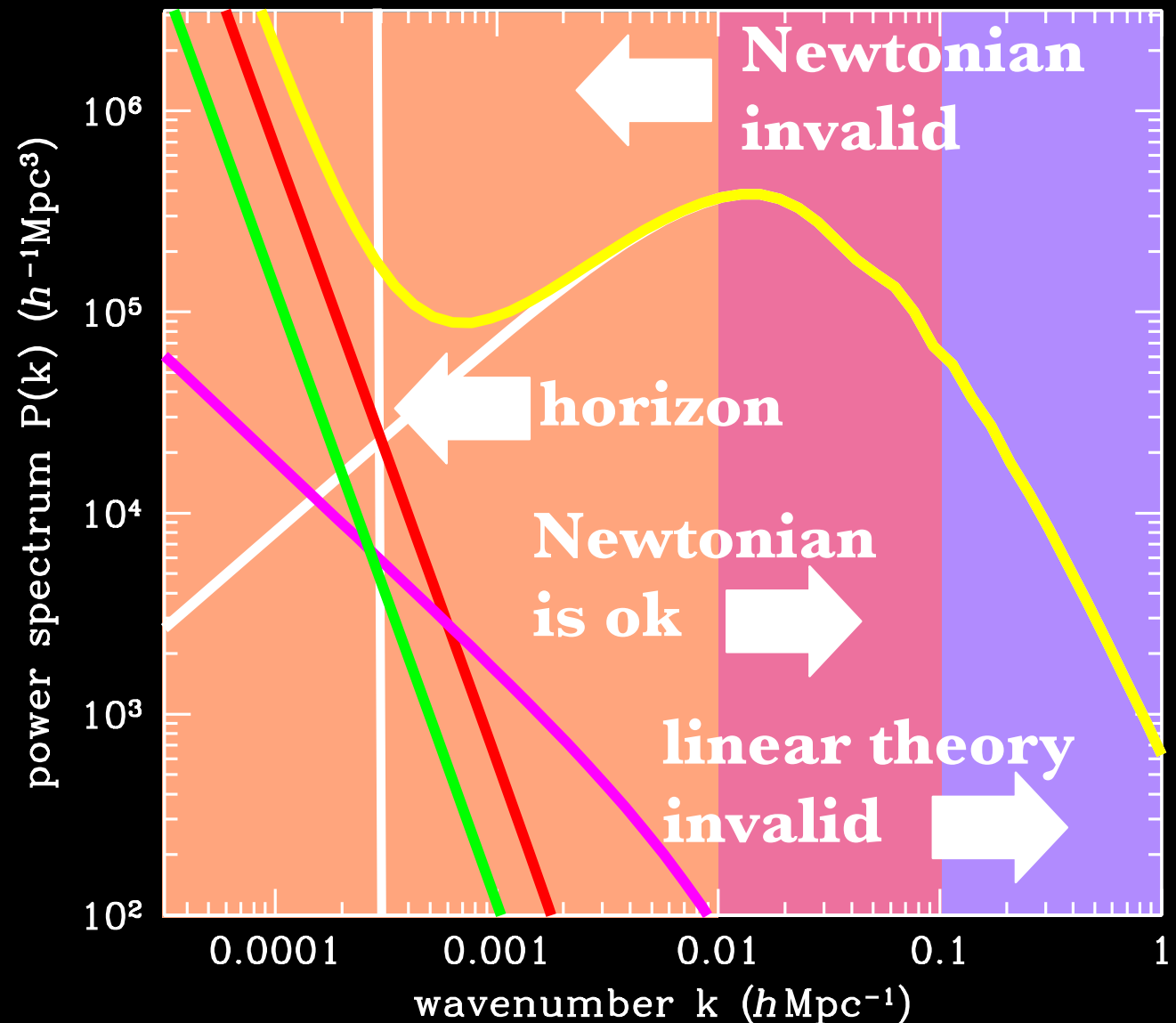
$$P_\eta(k)$$

- **Newtonian**

$$P_\delta(k)$$

$$P_\psi(k)$$

$$P_v(k)$$



# Scales of Interest

$z=6, 3, 1, 0.5$

- **Synchronous**  
(CMBFast, CAMB)

$$P_{\delta}(k)$$

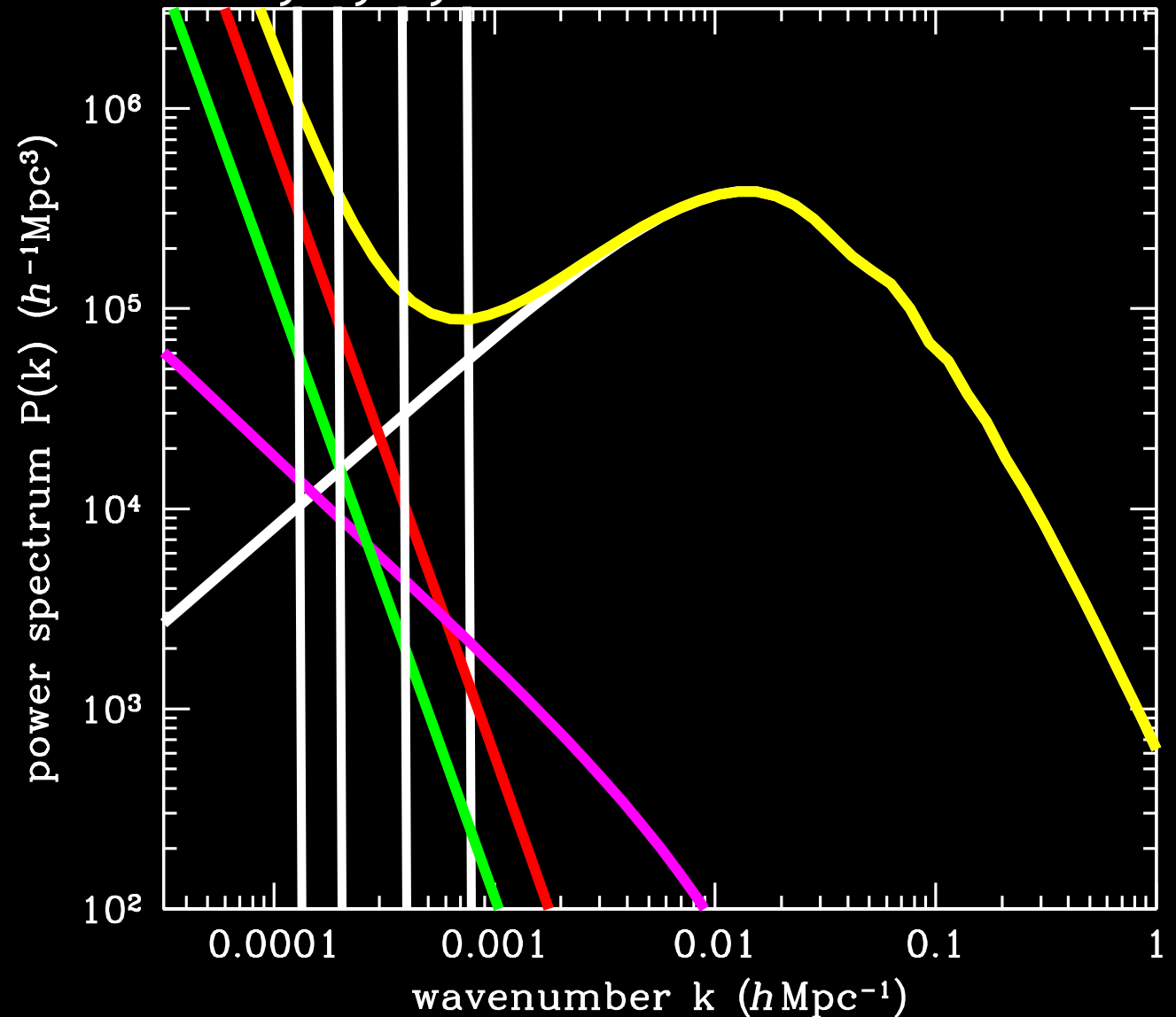
$$P_{\eta}(k)$$

- **Newtonian**

$$P_{\delta}(k)$$

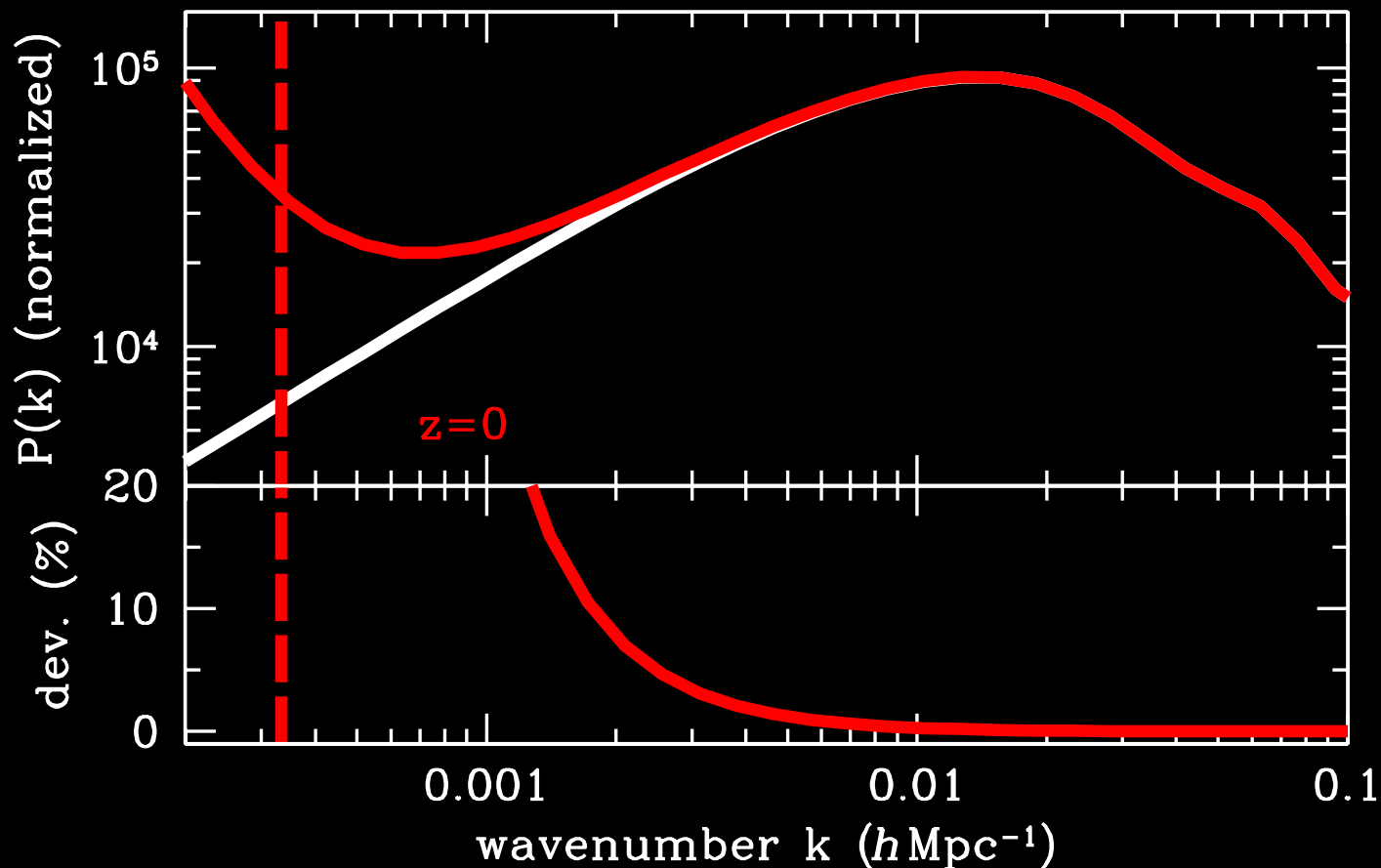
$$P_{\psi}(k)$$

$$P_v(k)$$



## Hubble Horizon

- horizon scale  $\sim 3 \text{ Gpc}/h$  (today),  $\sim 3 \text{ deg.}$  (recomb.)  
*relativistic effects are order one at horizon scale!*



## Summary

- *fully gauge-invariant general relativistic description*
- **standard method:** *gauge-dependent!*
  - galaxy clustering
  - gravitational lensing
- **in Newtonian limit:** *Do not worry!*
  - general relativity reduces to Newtonian

### **III. APPLICATIONS:**

*Impacts on Current Surveys – Why Bother?*

## Systematic Errors

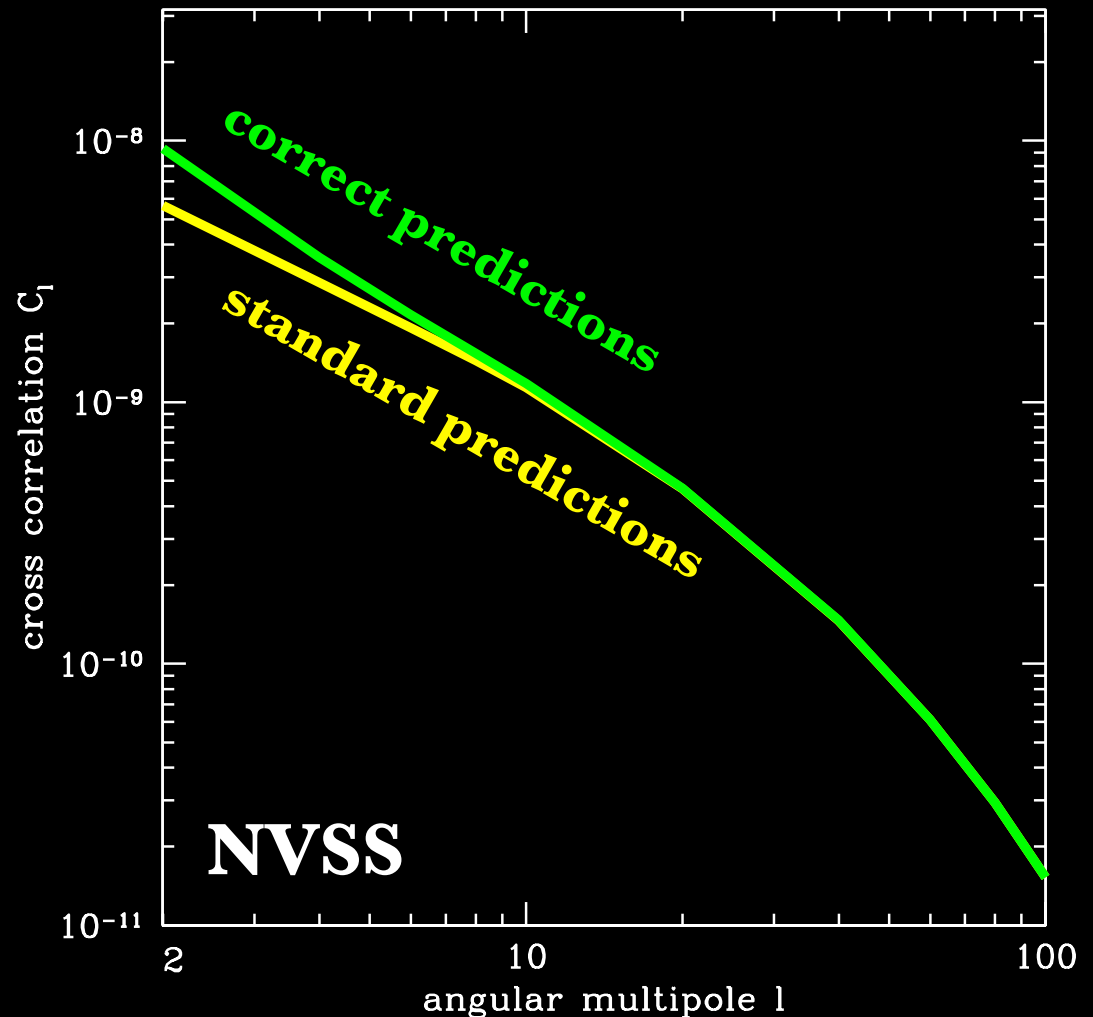
- **theoretical predictions:**

- **new cal.** (*correct*)
- **standard** (*incorrect*)

- **standard method:**

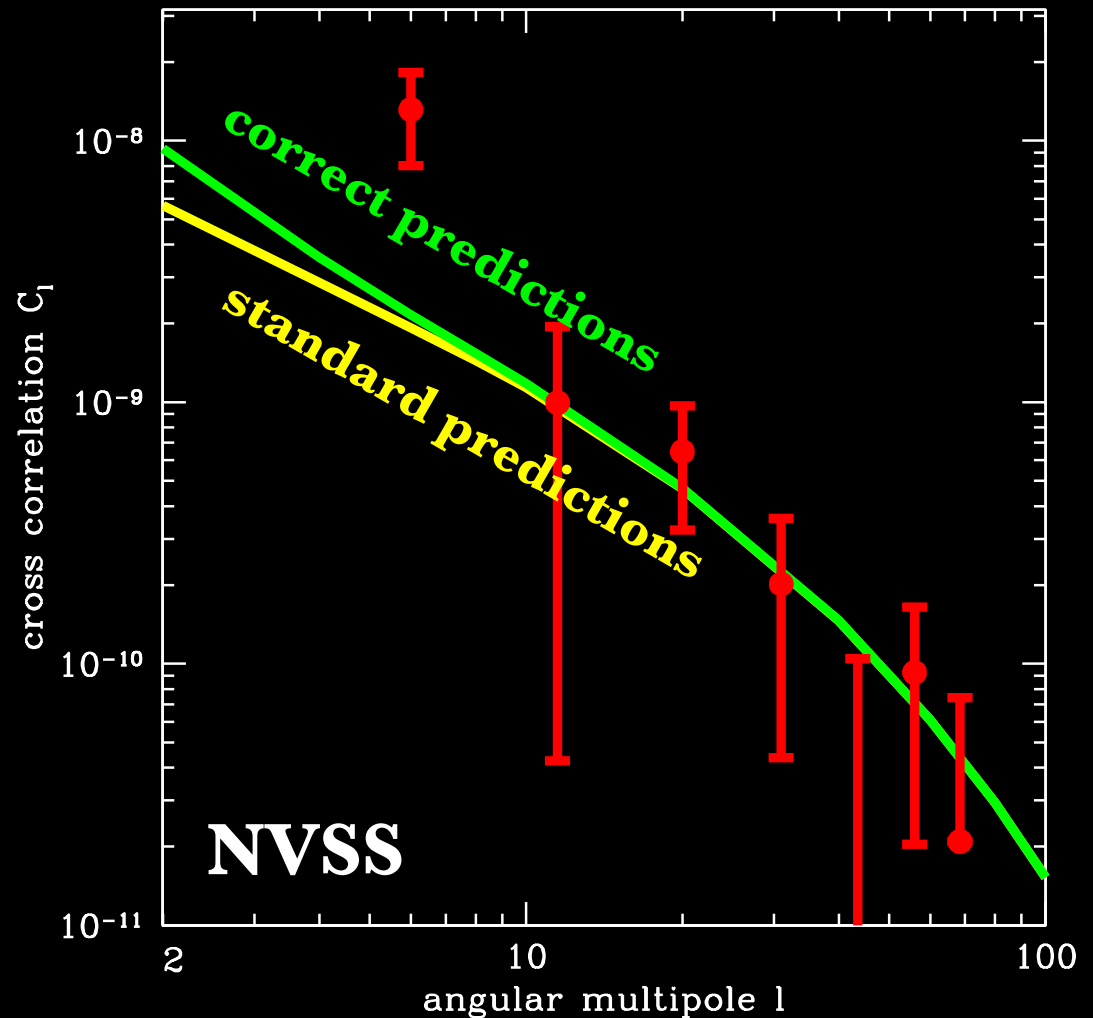
$$\delta_g = b \delta_m^{sync} + (5p - 2) \kappa$$

- *underestimate* the observed signals by *a factor two* at low multipoles



## Systematic Errors

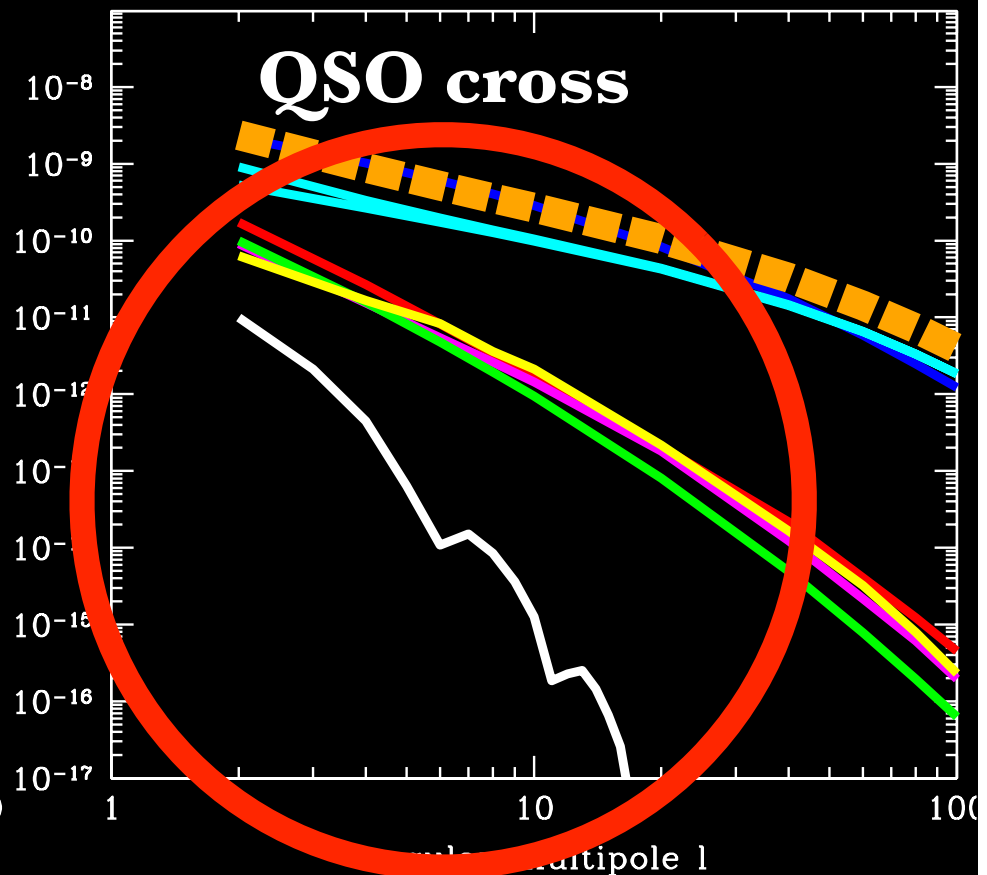
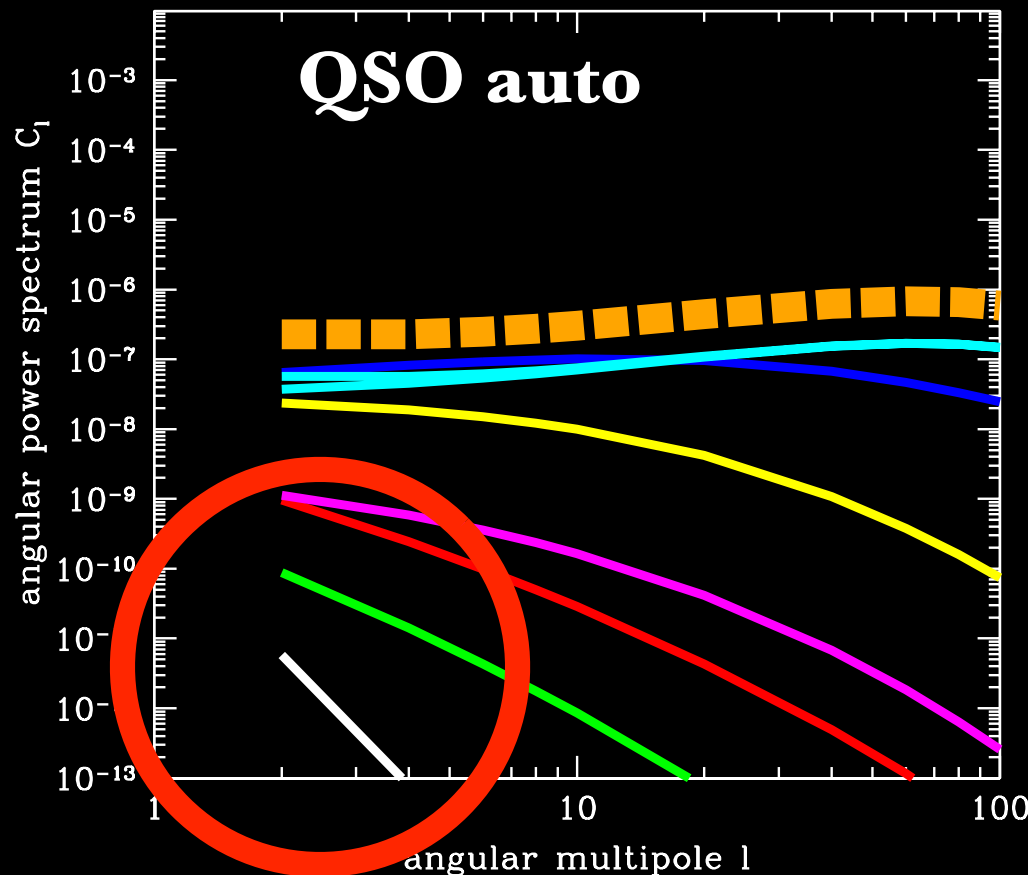
- theoretical predictions:
  - new cal. (*correct*)
  - standard (*incorrect*)
- 3.7- $\sigma$  detection, but *observed signal is larger by 2 at low multipoles* when all tracers are combined
- *anomalous large signal*



Ho, Seljak et al., PRD, 2008

# Primordial Gravity Waves

- integrated Sachs-Wolfe effect *for* gravity waves
  - tensor-to-scalar ratio  $r=0.1$

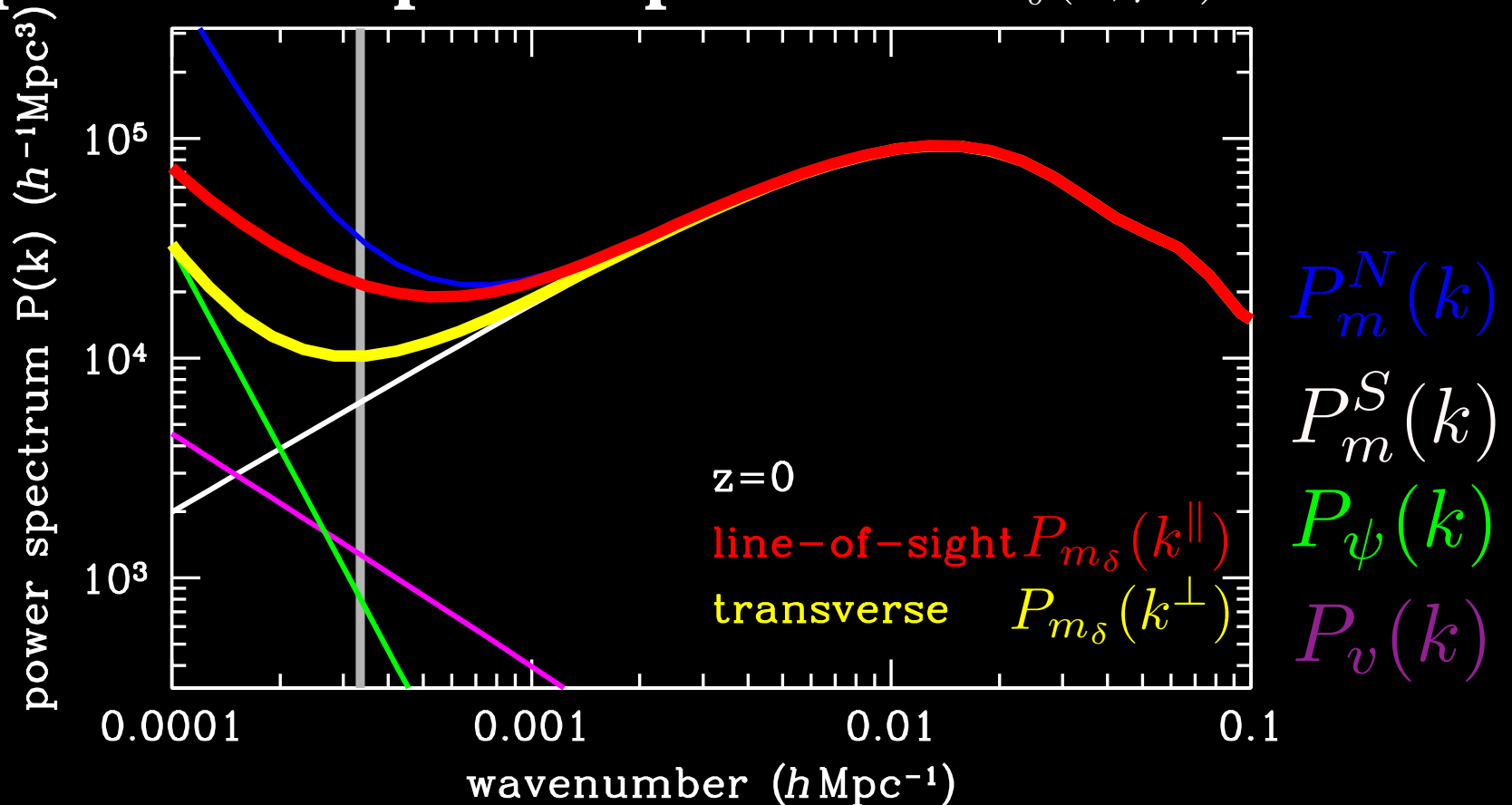


# Galaxy Power Spectrum

- **matter fluctuation:**  $\rho_m = \bar{\rho}_m(t)[1 + \delta_m] = \bar{\rho}_m(z_{\text{obs}})[1 + m_\delta]$ 
  - gauge-dependent  $\delta_m$   $1 + z_{\text{obs}} = (1 + z)(1 + \delta z)$
  - gauge-invariant, *observable*  $m_\delta = \delta_m - 3 \delta z$
  - Bardeen's gauge-invariant  $\epsilon_m, \epsilon_g$  ( $\epsilon_m \neq \epsilon_g \neq m_\delta$ )  
*correct “real-space” matter fluctuation!*
- **time slicing** (hypersurface of simultaneity)
  - constant coordinate time for  $\delta_m, \bar{\rho}_m(t)$
  - observed redshift  $z_{\text{obs}}, \bar{\rho}_m(z_{\text{obs}})$
  - matter rest frame, zero-shear frame

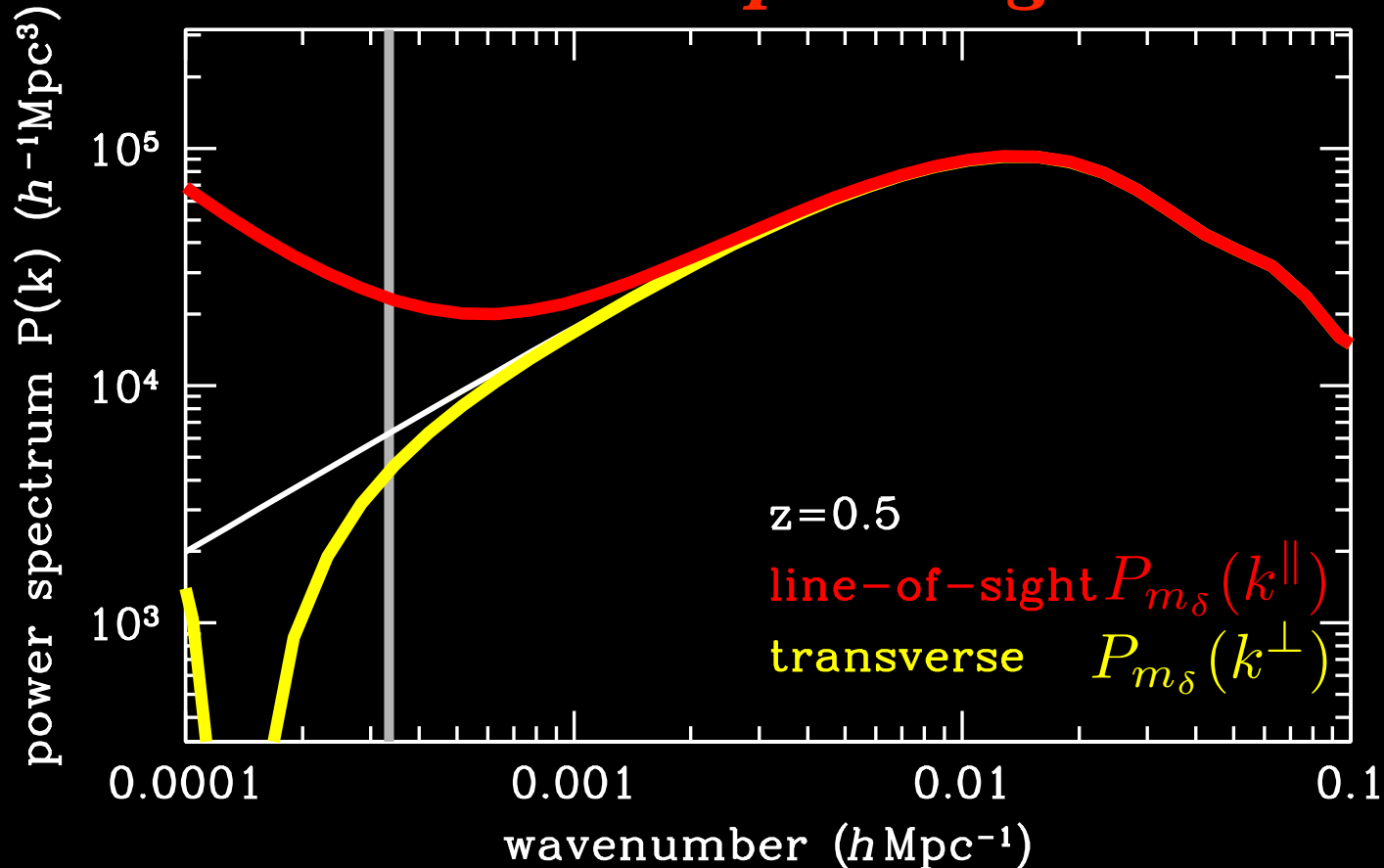
## “Real-Space” Power Spectrum

- *no longer isotropic*, neither  $P_m^S(k)$ , nor  $P_m^N(k)$
- **real-space matter power spectrum**  $P_{m_\delta}(k, \mu_k)$



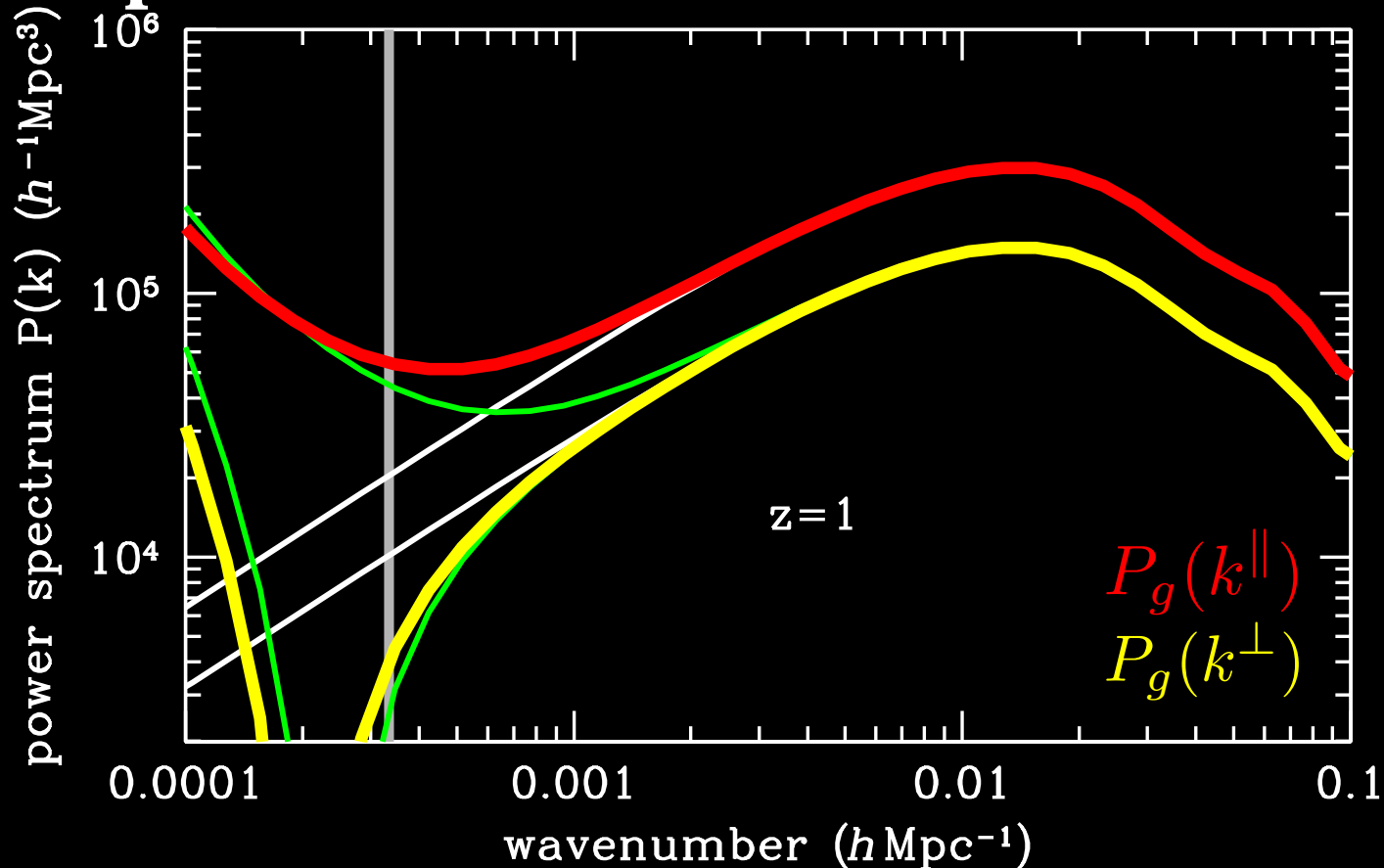
# “Real-Space” Power Spectrum

- appear as an underdense region on large scales
- *characteristic scale* and *shape change*



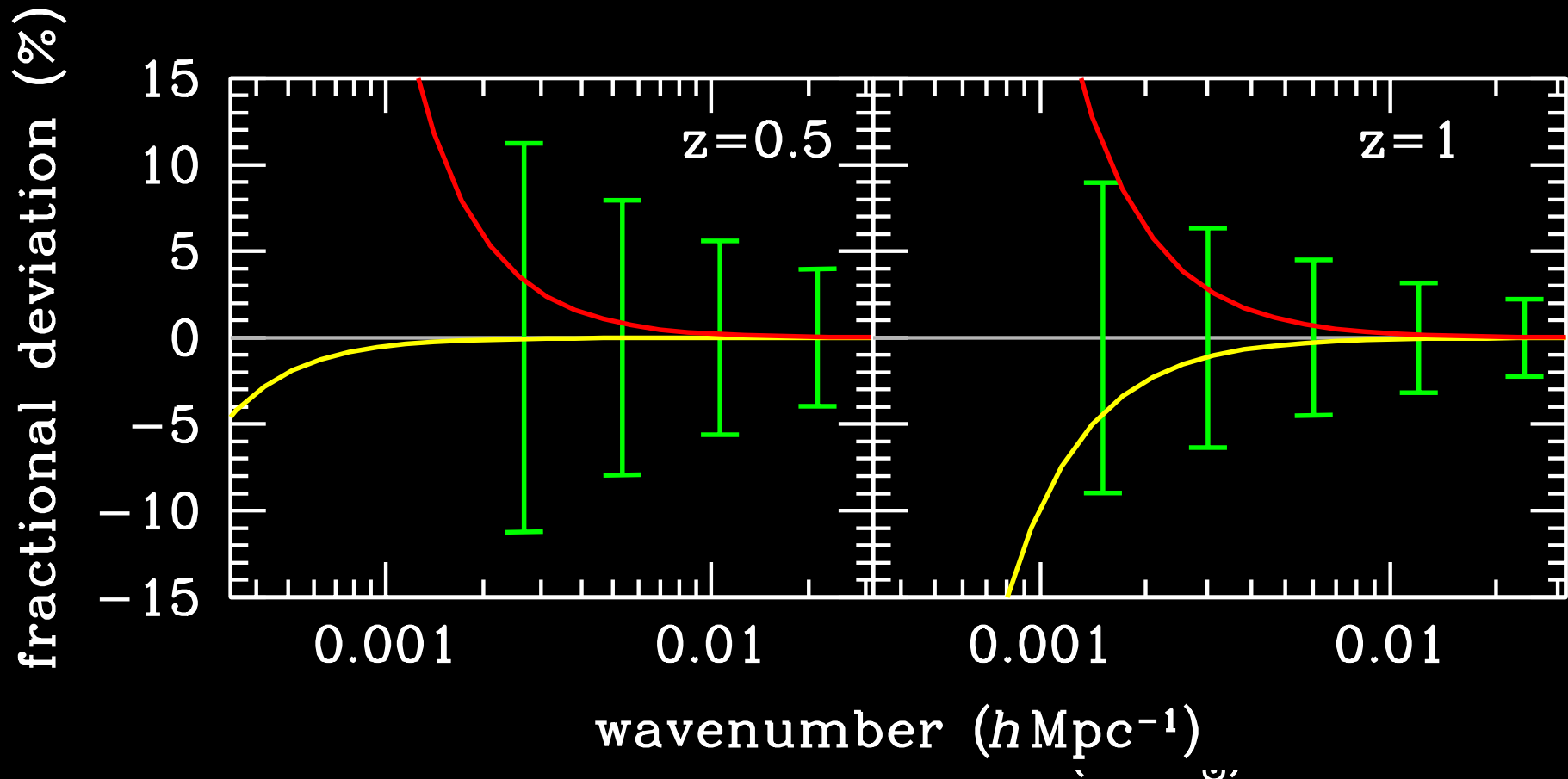
# Full Galaxy Power Spectrum

- largely similar to real-space matter power spectrum
- redshift-space distortion boost



## Systematic Errors

- **B**aryonic **O**scillation **S**pectroscopic **S**urvey (BOSS)



## **IV. SUMMARY AND PROSPECTS:**

*What do We Learn from This?*

## New Perspective on GC

- *new perspective* on  
galaxy clustering as a cosmological probe
- **unified treatment:**
  - redshift-space distortion, magnification bias, ...
- **subtle gauge issues:** *observables*
  - redshift, magnification, galaxy bias
- **general relativistic description:**
  - gauge-invariant formalism
  - *tensor* as well as scalar contributions

# Relativistic Effects

- with optimal weighting and multi-tracer method  
GR effect in galaxy clustering *reality & measurable*
- **future work:**
  - constraints on *modified gravity* on horizon scales  
complementary to the Solar system constraints
  - constraints on *dark energy* models with  
superhorizon perturbations

# New Perspective on Cosmology

- galaxy clustering as a worked example
- *many* theoretical predictions in cosmology
  - gauge-dependent, unobservable
  - how to remedy other probes?
  - *follow observational procedure*

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**JAIYUL YOO**

**INSTITUTE for THEORETICAL PHYSICS, UNIVERSITY of ZÜRICH  
LAWRENCE BERKELEY LABORATORY, UNIVERSITY of CALIFORNIA, BERKELEY**

**Korea Institute for Advanced Study, June, 28, 2011**